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Relativistic Dynamics

* mass: The mass m of an object is defined in Rest Frame (RF) - 4-scalar

* world line is defined by 4-vector $x = (\vec{x}(t), ct)$

* proper time is measured in RF

dt_0 : $dx_0 = x_0(t+dt_0) - x_0(t) = (0, 0, 0, c dt_0)$
 moving frame: $dx = (v dt, c dt)$

Since $dx = dx_0 \rightarrow -c^2 dt_0^2 = (v^2 - c^2) dt^2 \rightarrow dt_0 = dt/\gamma$

dt_0 is the proper time interval - 4-scalar.

* 4-velocity $u = \frac{dx}{dt_0} = \left(\frac{dx}{dt_0}, c \frac{dt}{dt_0} \right)$ - 4-vector.

replace $dt_0 \rightarrow dt/\gamma \rightarrow u = \gamma \left(\frac{dx}{dt}, c \right)$ - 4-vector.

the spatial part of u is not velocity v
 it becomes so when $v \ll c$

* Relativistic momentum

- should produce $p = mv$ in classical limit

- should be conserved for isolated system in all frames

if classical momentum $\sum p_i = \sum m_i v_i$ is conserved in S , it may not be conserved in S'

- relativistic momentum: $p = m u = (\gamma m v, \gamma m c)$

m - 4-scalar u - 4-vector $\rightarrow QR \rightarrow P$ is 4-vector.

conservation law: $\sum P_{in} = \sum P_{out}$

includes both conservation of 3D momentum and energy. $E = \gamma m c^2 = P_4 c$

$P = (\vec{P}, E/c)$ - momentum energy 4-vector.

classical limit: $E \approx m c^2 + \frac{1}{2} m v^2$

classical conservation of energy:

$\sum_i E_{in}^i = \sum_i E_{fin}^i \rightarrow \sum_{in} (m_i c^2 + \frac{1}{2} m_i v_i^2) = C_{in} + T_{in} = C_{fin} + T_{fin}$

conservation of mass $C_{in} = C_{fin} \rightarrow$ kinetic energy T is conserved.

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inelastic collision: $T_{in} \neq T_{fin}$.

→ $C_{in} \neq C_{fin}$ mass is not conserved if.
 We believe in conservation of energy $E_{in} = E_{fin}$.

$$\Delta m c^2 = -\Delta T \rightarrow \Delta M = -\Delta T/c^2 \quad \text{usually small!}$$

rest energy: $E = \gamma m c^2 \quad \gamma \rightarrow 1 \quad \boxed{E = m c^2}$

annihilation of $e^+ e^- \rightarrow 2\gamma$ (radiation)
 mass \rightarrow radiation.

$$\bar{E} = m c^2 + T \quad \text{- definition of kinetic energy}$$

$$\gamma m c^2 = m c^2 + T \quad T = m c^2 (\gamma - 1)$$

$$\beta E = p c; \quad p \cdot p = -m c^2; \quad \bar{E}^2 = (m c^2)^2 + (p c)^2$$

Collisions: Example 15.9 $m \xrightarrow{v} M \rightarrow M \xrightarrow{u=?}$

$$P_{in}^2 = (P_m + P_M)^2 = P_m^2 + P_M^2 + 2 P_m P_M = -m c^2 - M c^2 + 2 P_m P_M$$

$$P_m = (p, \gamma m c) \quad (P_M = 0, M c) \rightarrow 2 P_m P_M = (0 - \frac{E_m}{c} \cdot M c)^2$$

$$P_{in}^2 = -m c^2 - M c^2 - 2 E_m M$$

$$P_{fin} = (0, M_{tot}) \quad P_{fin}^2 = -M_{tot}^2 c^2$$

$$M_{tot} = \sqrt{m^2 + M^2 + 2 \frac{E_m}{c^2} M}$$

classical limit $M_{tot} = (m+M)$
 $E_m = m c^2$

$$v = \frac{\bar{P}_{fin} c^2}{E_{fin}} = \frac{\bar{p} c^2}{E_m + M c^2} =$$

CM frame: $\vec{P} = \sum \vec{p} = 0 \quad p \cdot p = (0, 0, 0, P_4)$