Paul McNamara on behalf

2Sa science

of the LPF Team

LISA Symposium X

University of Florida, May 2014

Monday 19 May 14

lisa pathfinder

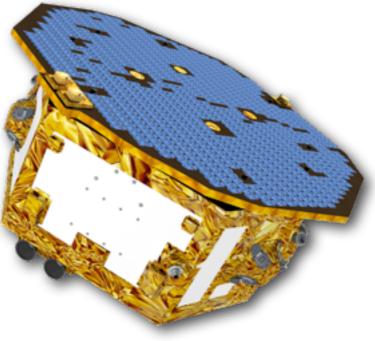
Introduction



- LISA Pathfinder (LPF) is a technology verification mission for future spaceborne gravitational wave detectors
 - LPF was approved by ESA to demonstrate the concept of low frequency gravitational wave detection in space

The LPF mission is a stepping stone in the development of a spaceborne GW mission

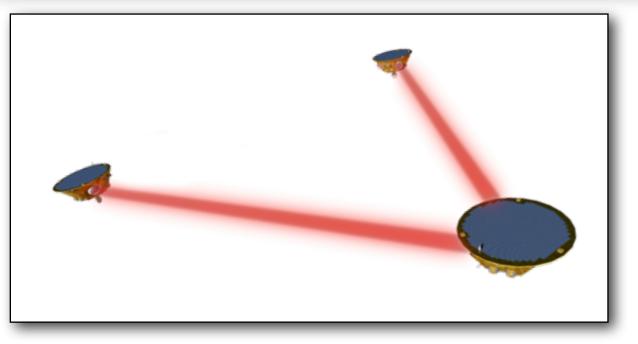
- LISA Pathfinder provides us with:
 - A better understanding of the physics of the forces acting on a free-falling test mass
 - Industrial experience in the development, manufacture, and testing of technologies required for GW detection
 - Data analysis algorithms and tools dedicated to the analysis of the system as a whole
 - Essential experience in the commissioning of a LISAlike mission





LISA Pathfinder





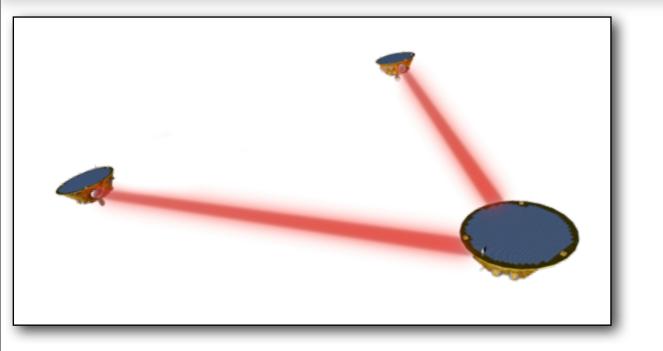
(e)LISA:

- 3 spacecraft, separated by ~1million km

- Role of each spacecraft is to protect the fiducial test masses from external forces

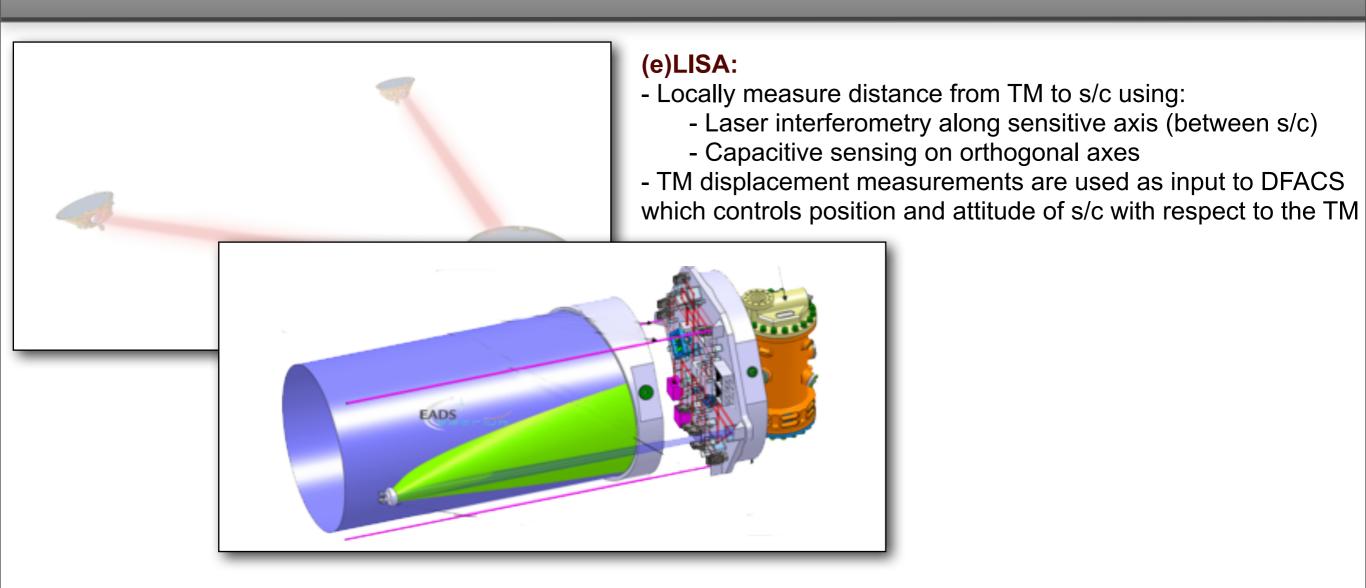






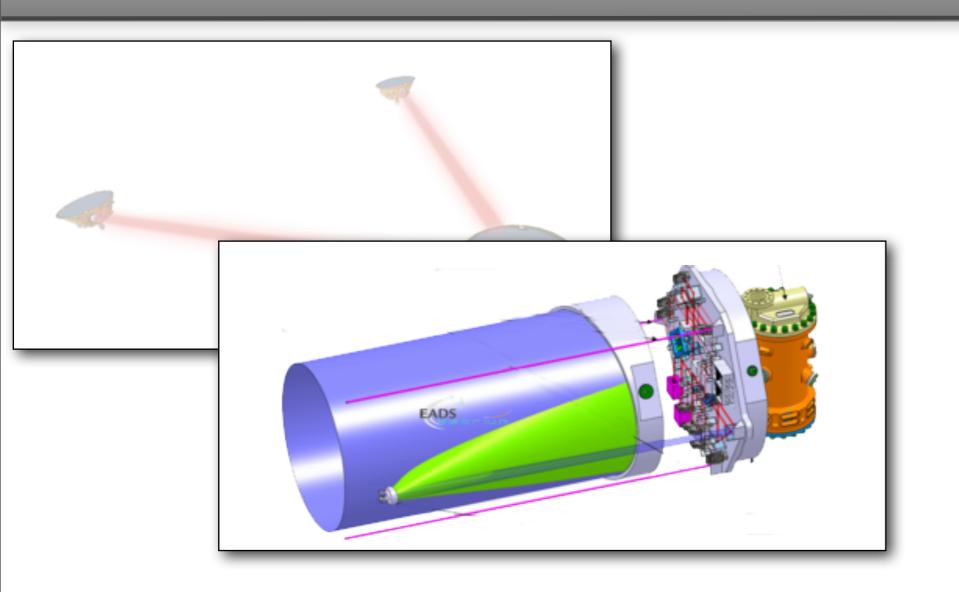




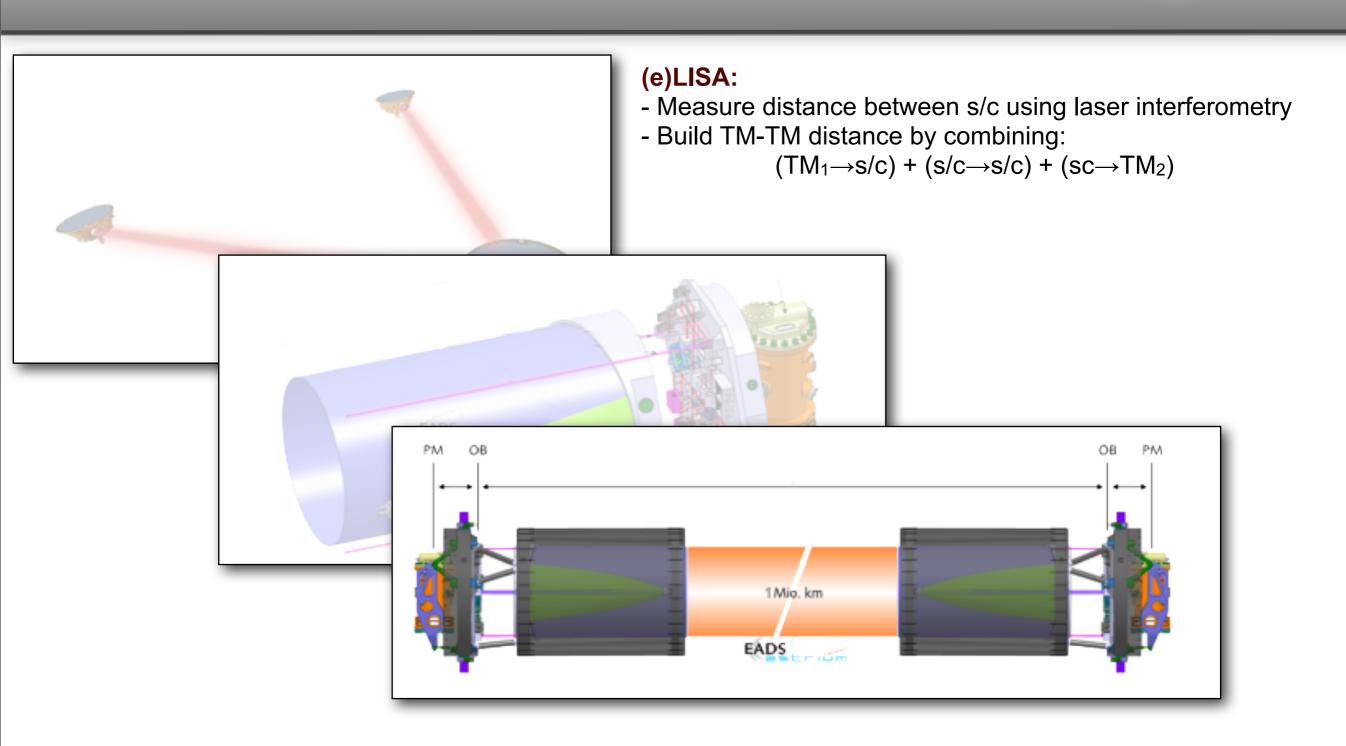








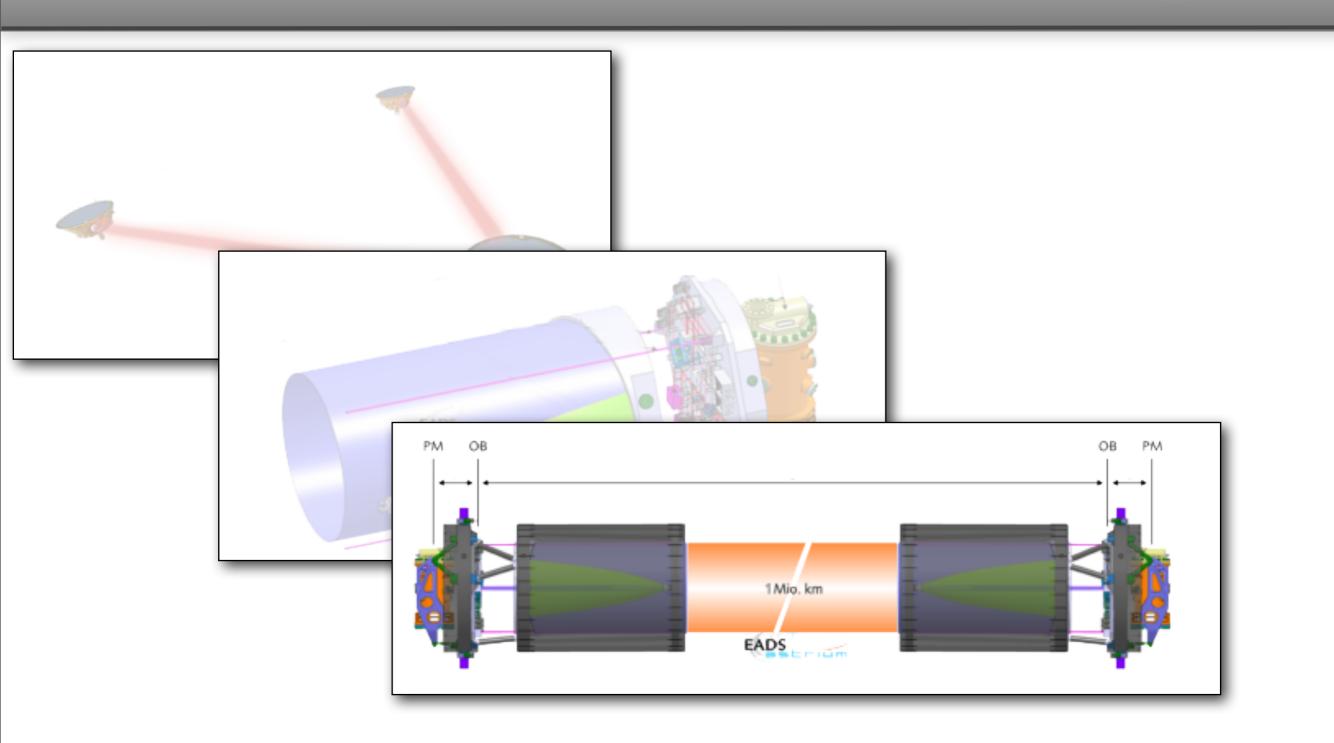






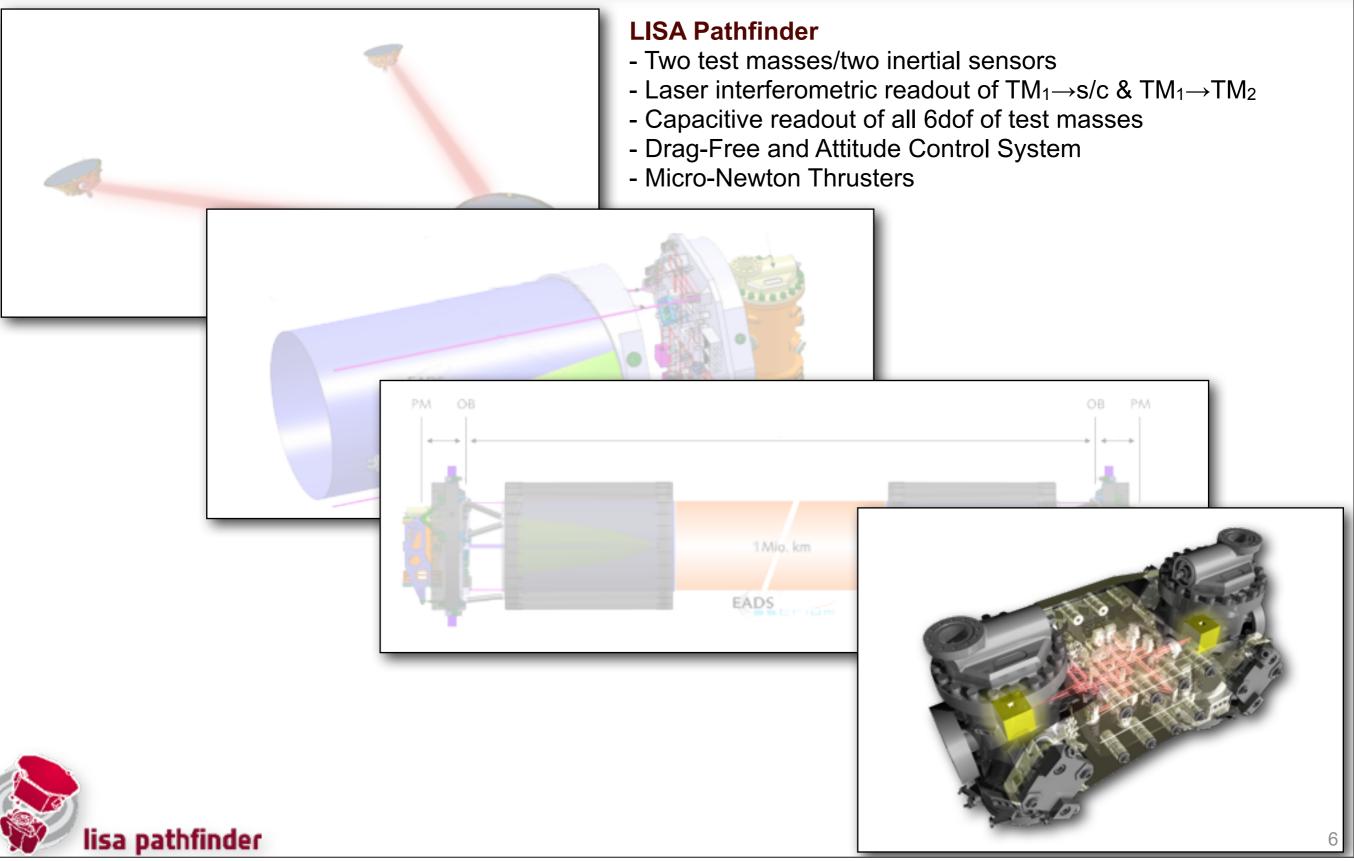
esa













- > LISA Pathfinder has two primary performance requirements:
 - Differential Acceleration Noise between free-floating test masses

$$S_{a}^{\frac{1}{2}}(f) \le 3 \times 10^{-14} \left[1 + \left(\frac{f}{3 \,\mathrm{mHz}}\right)^{4} \right]^{\frac{1}{2}} \,\mathrm{ms}^{-2}/\sqrt{\mathrm{Hz}}$$

- Displacement sensing noise

$$S_{oms}^{\frac{1}{2}}(f) \le 9.1 \times 10^{-12} \left[1 + \left(\frac{3 \,\mathrm{mHz}}{f}\right)^4 \right]^{\frac{1}{2}} \,\mathrm{m/\sqrt{Hz}}$$



Primary Performance Requirements



> LISA Pathfinder has two primary performance requirements:

 Differential Acceleration Noise between free-floating test Bill Weber: "LPF: achieving and measuring sub-femto-g free fall for

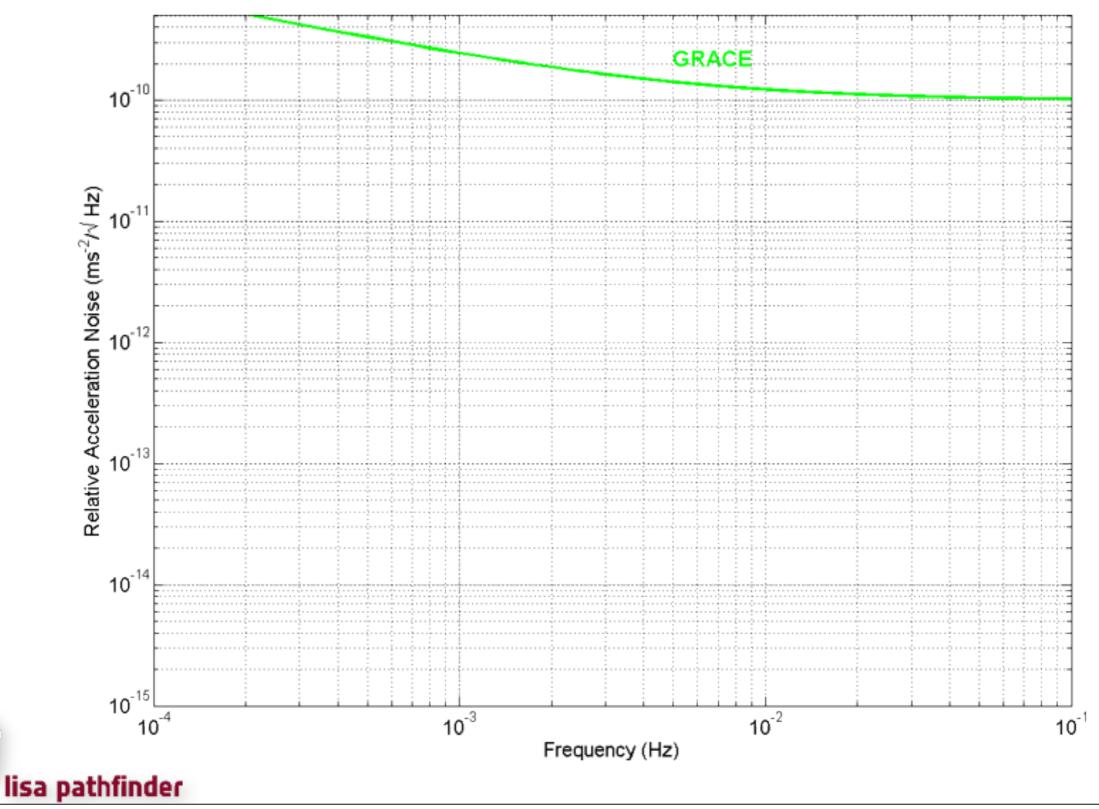
gravitational wave astrophysics"

Displacement sensing noise

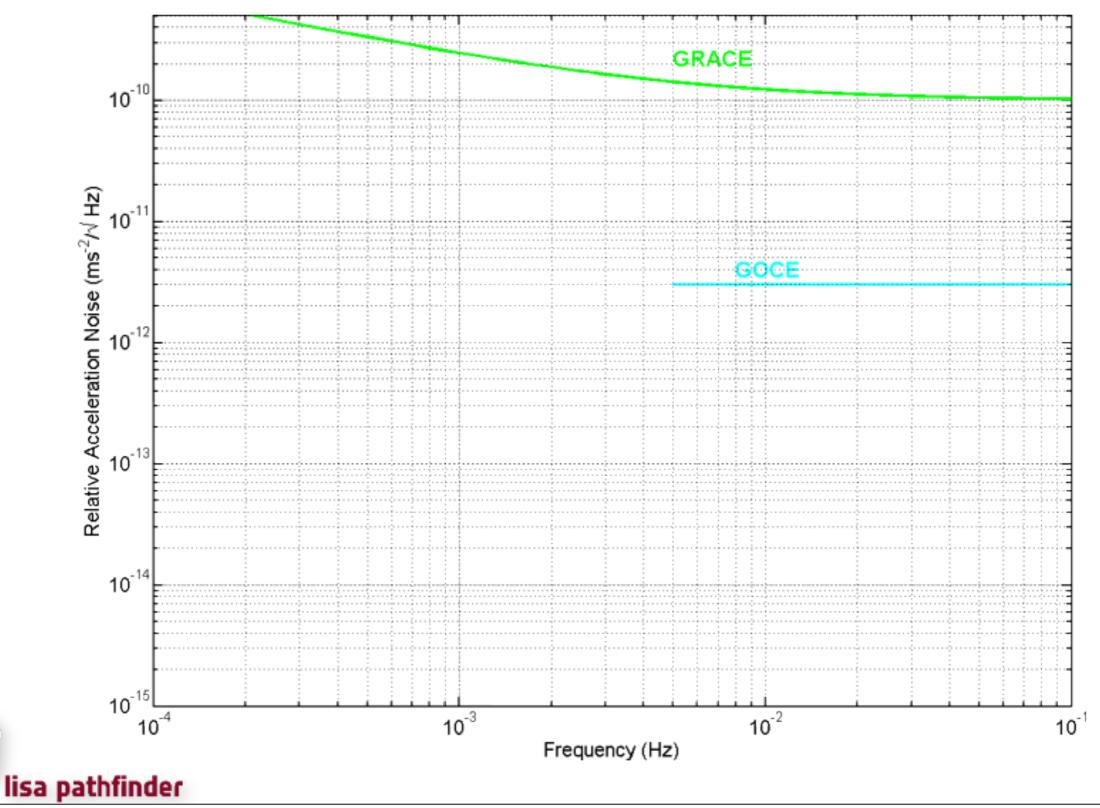




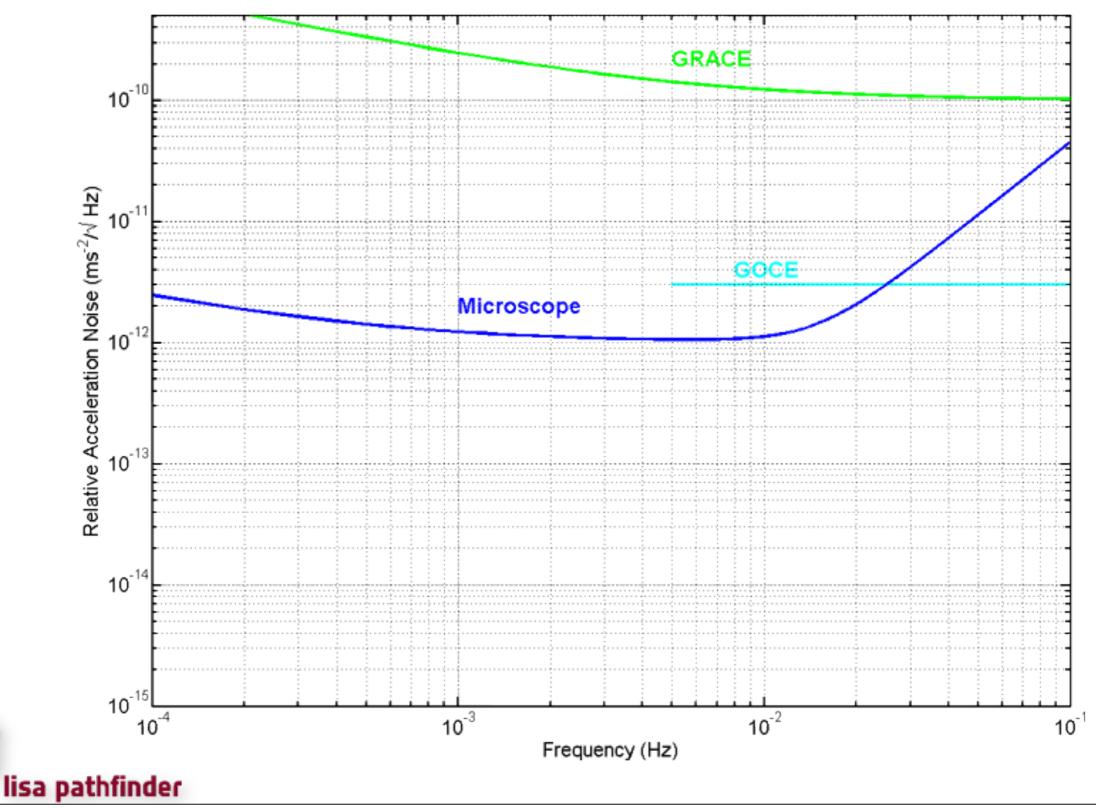




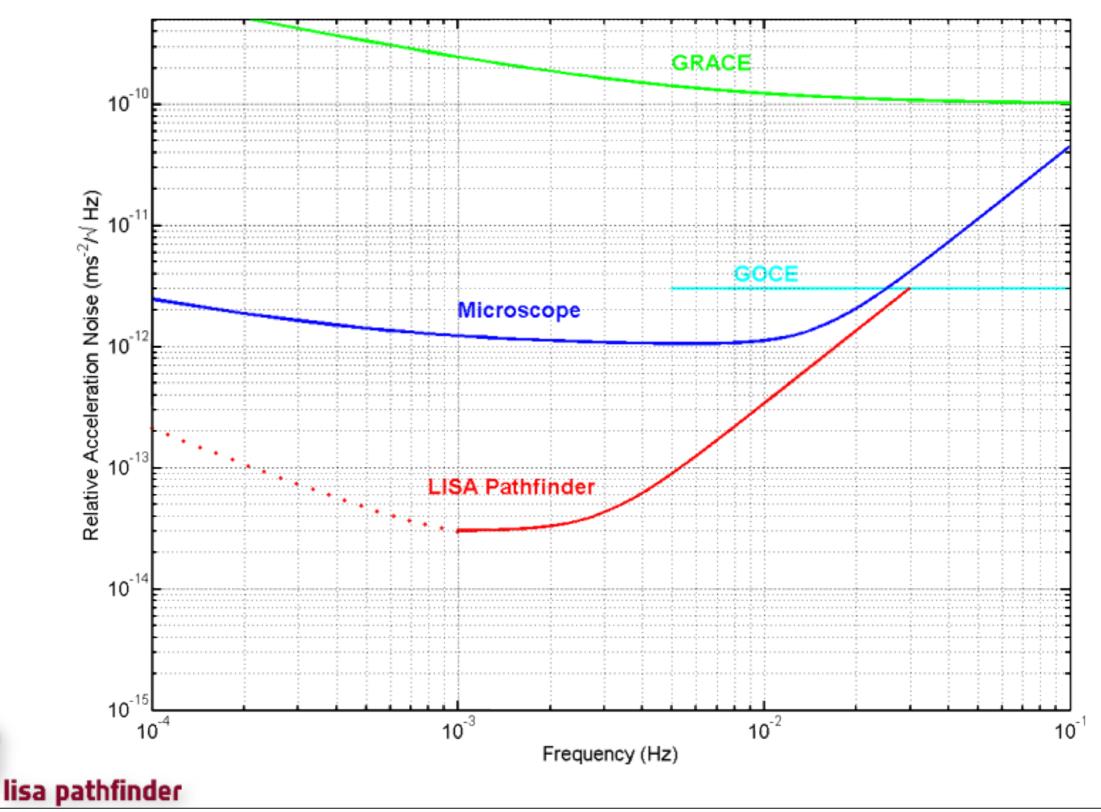




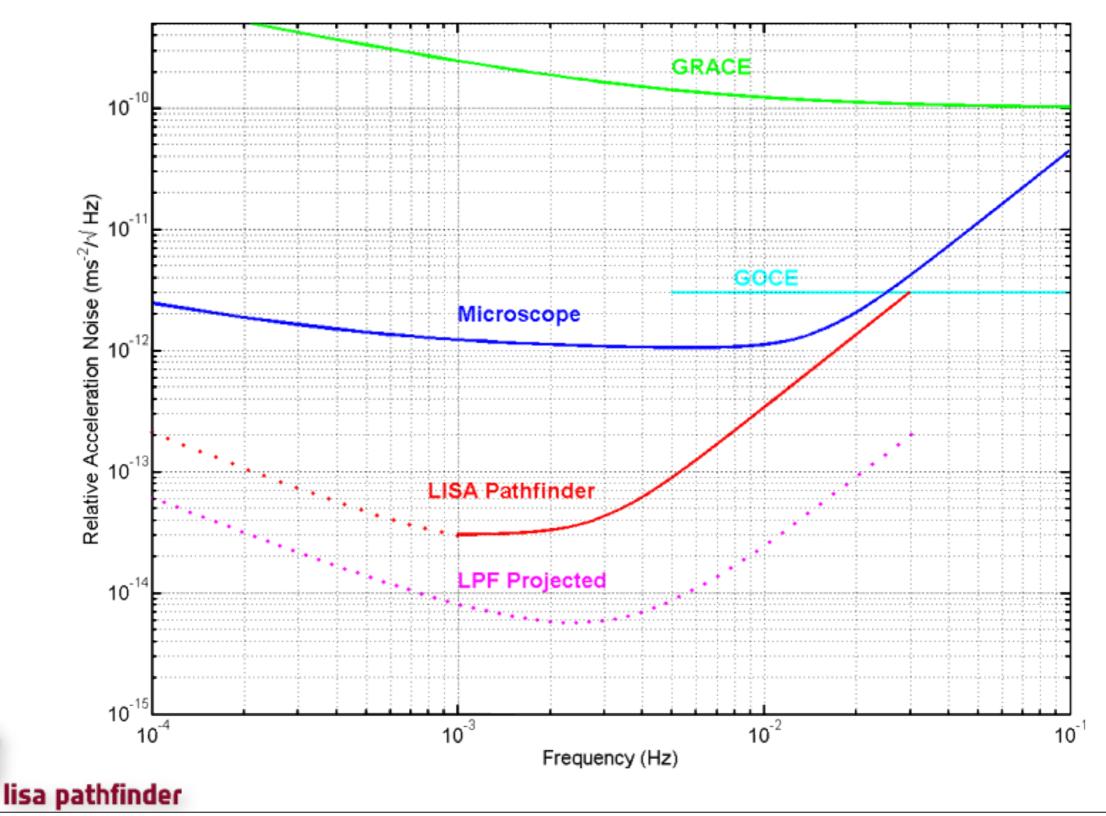




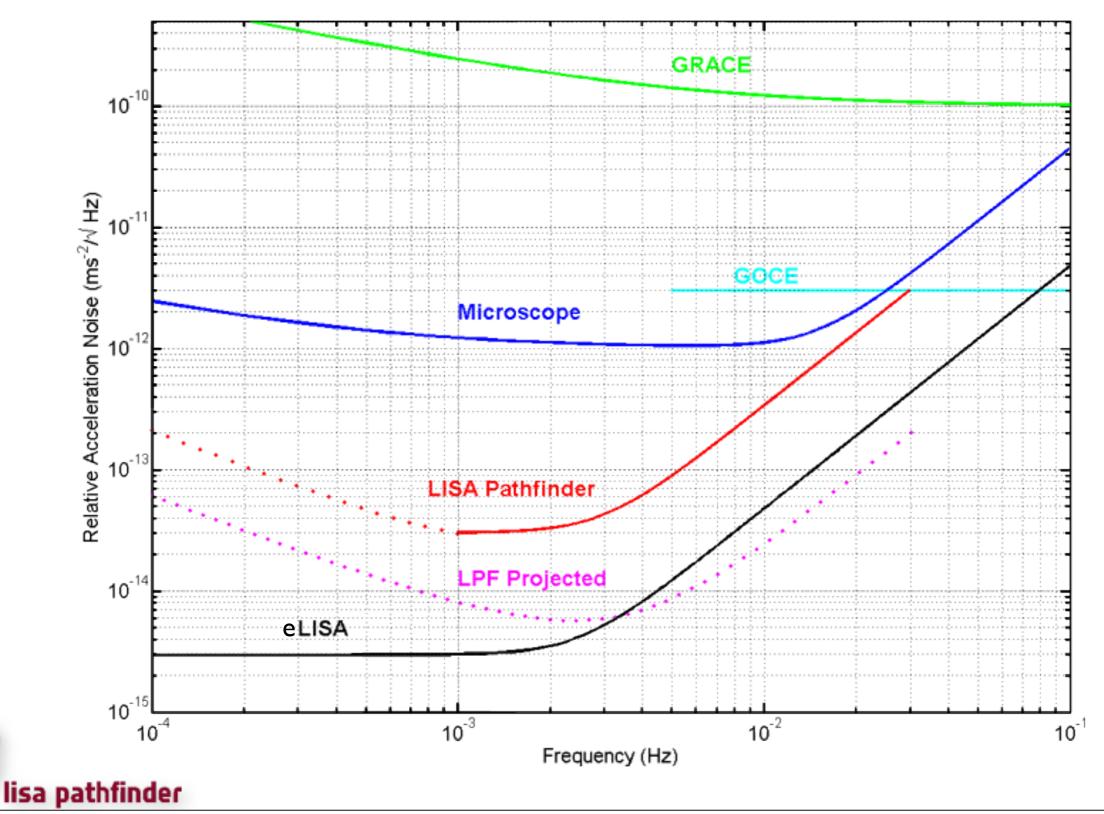




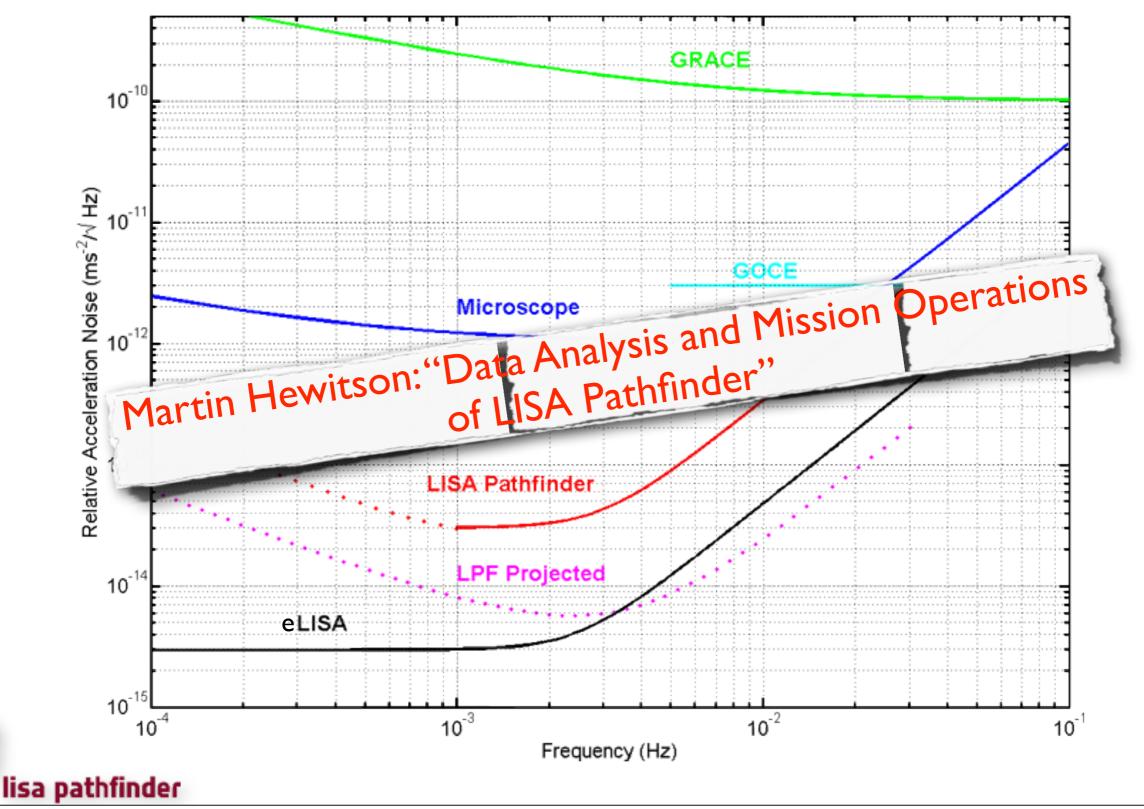






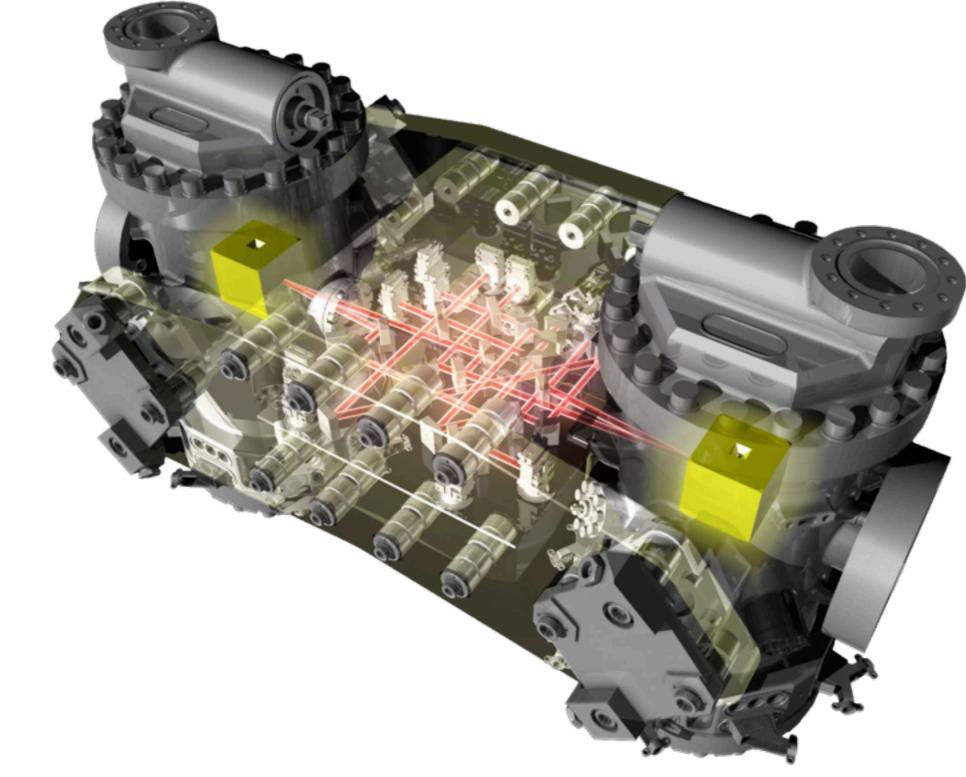






LISA Technology Package









Procurement and manufacture of the LTP funded by European member states and ESA



France:

Laser modulator

Germany: PI, LTP Architect (Astrium), Laser

Italy:

PI, Inertial Sensor (ISS)

Netherlands:

ISS SCOE

Spain:

Data Diagnostics System, Data Management Unit Switzerland:

ISS Front End Electronics United Kingdom:

Optical Bench, Phase-meter, Charge Management

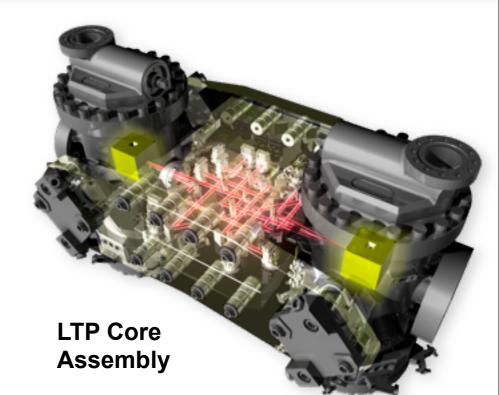
LISA Technology Package

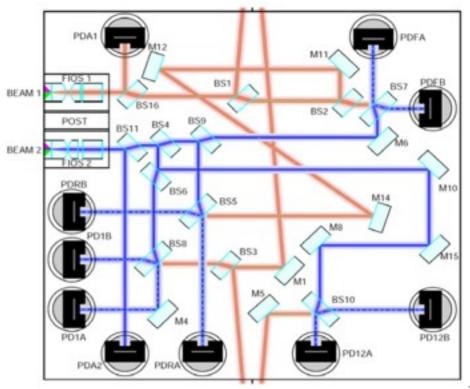


- The LISA Technology Package (LTP) is the European instrument payload on LPF
- Two Au:Pt test masses housed in separate vacuum enclosures
- Relative position of test masses read-out by
 - Heterodyne laser interferometry on sensitive axis
 - Capacitive sensing on all degrees of freedom
- Four interferometers on ultra-low expansion optical bench

interferometer

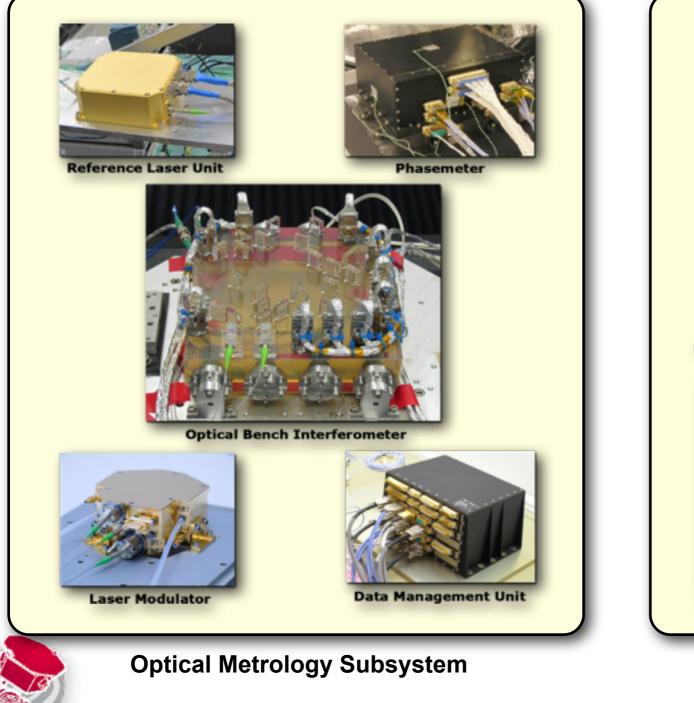
- x1, x2-x1, Frequency noise, Reference

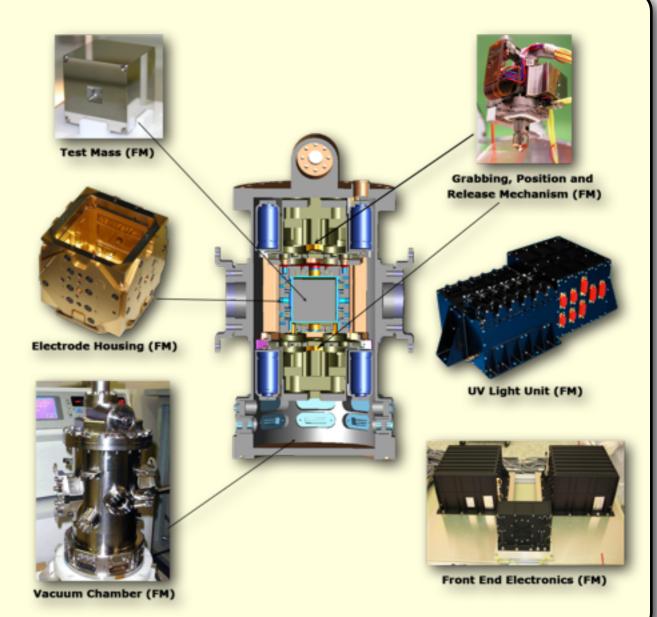






The LTP main subsystems



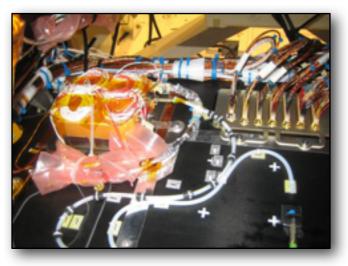


Inertial Sensor Subsystem

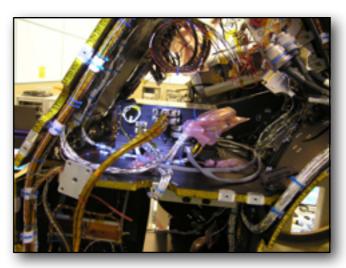


All electronic boxes have been integrated to the spacecraft

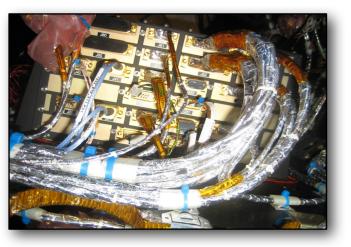
 Only the microthrusters and LTP Core Assembly (interferometer and inertial sensors) are not yet delivered



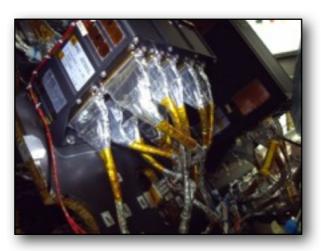
Reference Laser Unit



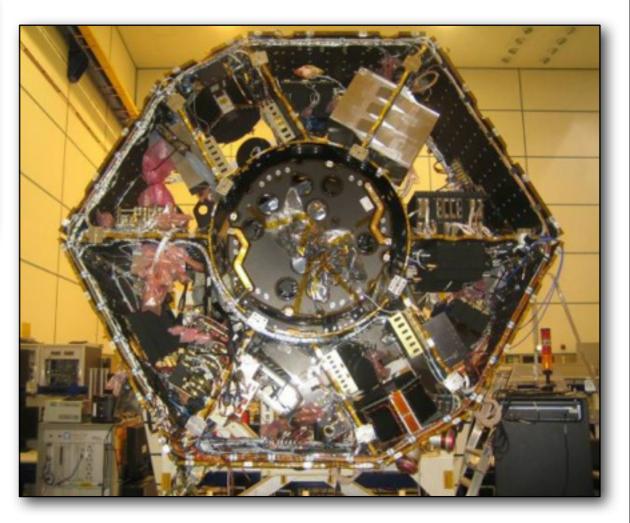
Phasemeter



Payload Computer



ISS Front-End Electronics



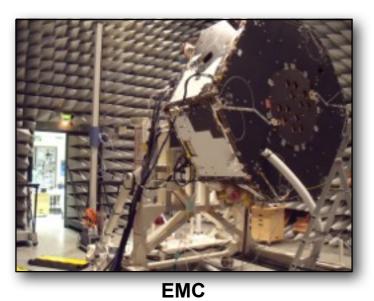
Integrated Sciencecraft. All photos courtesy of Airbus Defence and Space

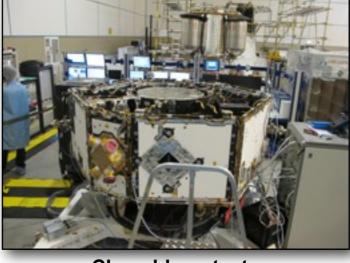


- Several system environmental tests have already taken place
 - Further system tests are on hold pending the delivery of the LTP Core Assembly



Vibration/shock tests





Closed-loop tests



On-Station Thermal Test





- Several system environmental tests have already taken place
 - Further system tests are on hold pending the delivery of the LTP Core Assembly

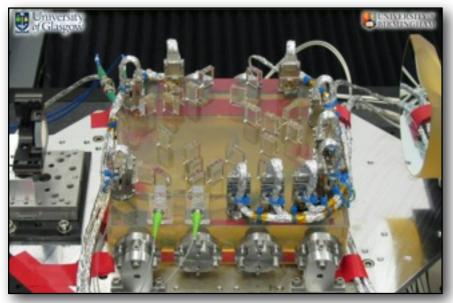


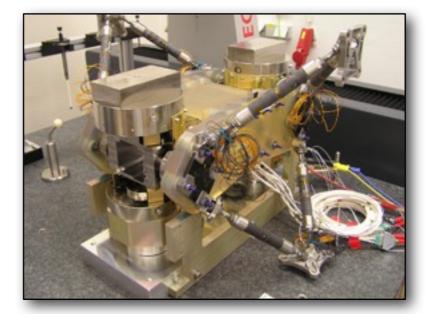


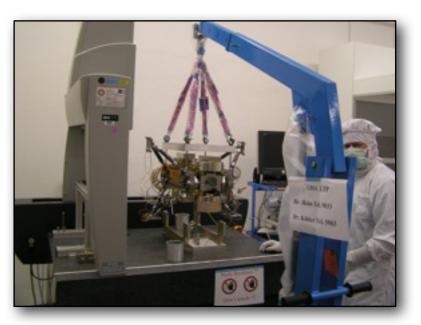
During the On-Station Thermal Test, we integrated a Thermal Optical Qualification Model (TOQM) in lieu of the LTP core assembly

 The TOQM consists of the flight optical bench with flight mounting hardware, and thermal dummies in place of the inertial sensors

The measured thermal and interferometric performance meets (and significantly exceeds) all requirements.

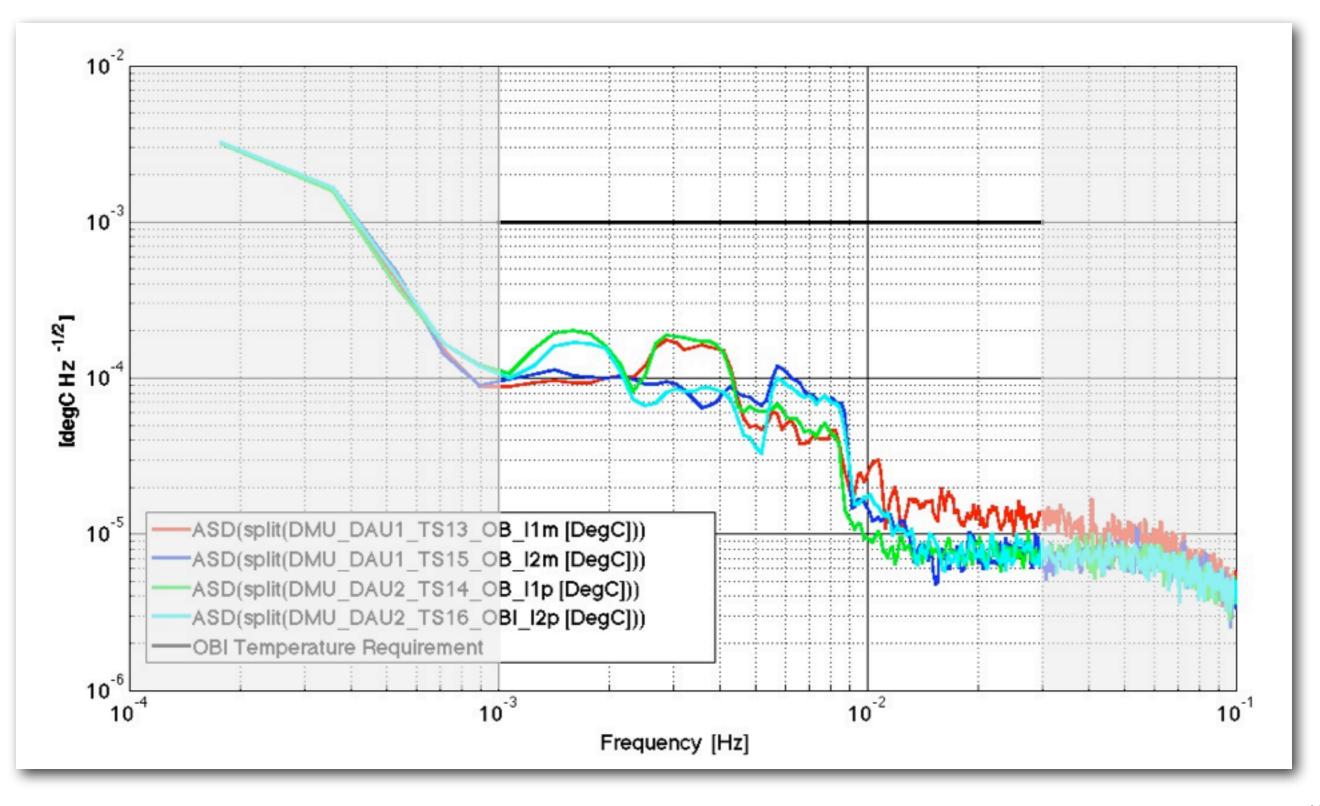








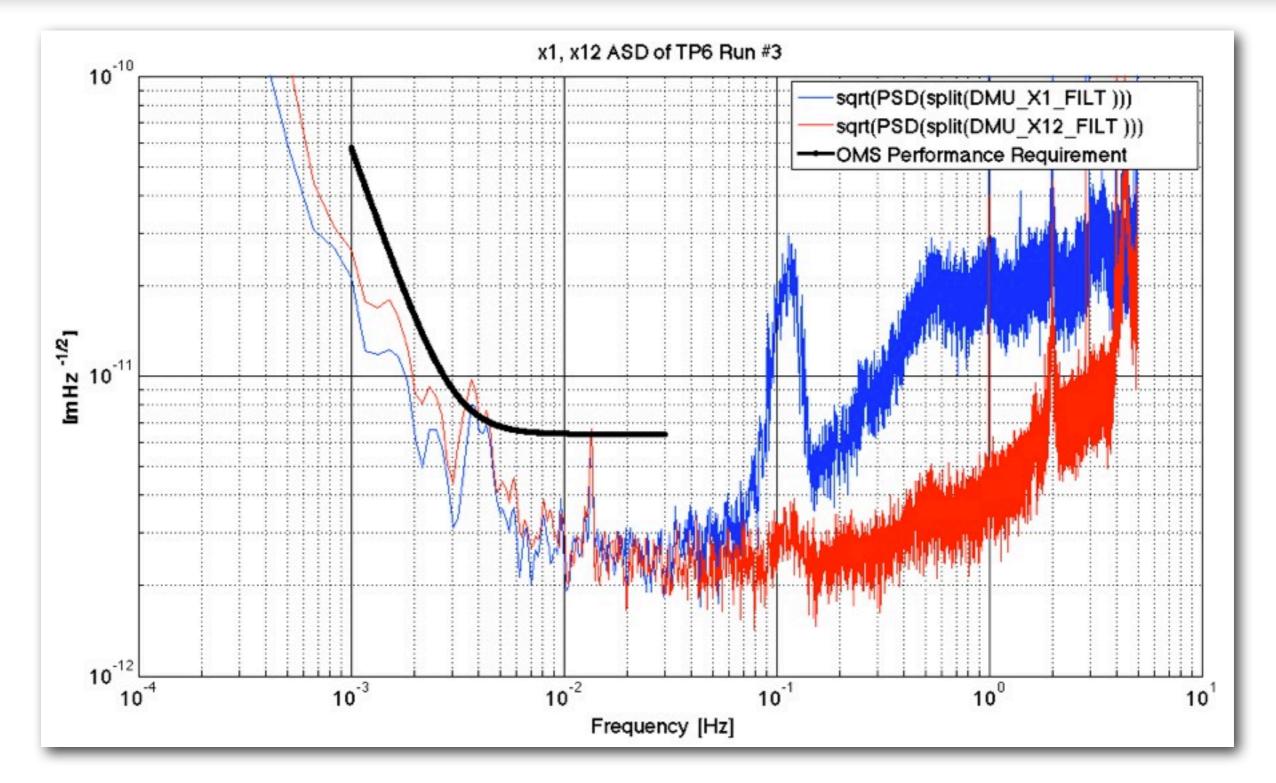
OSTT Thermal Performance



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OSTT OMS performance

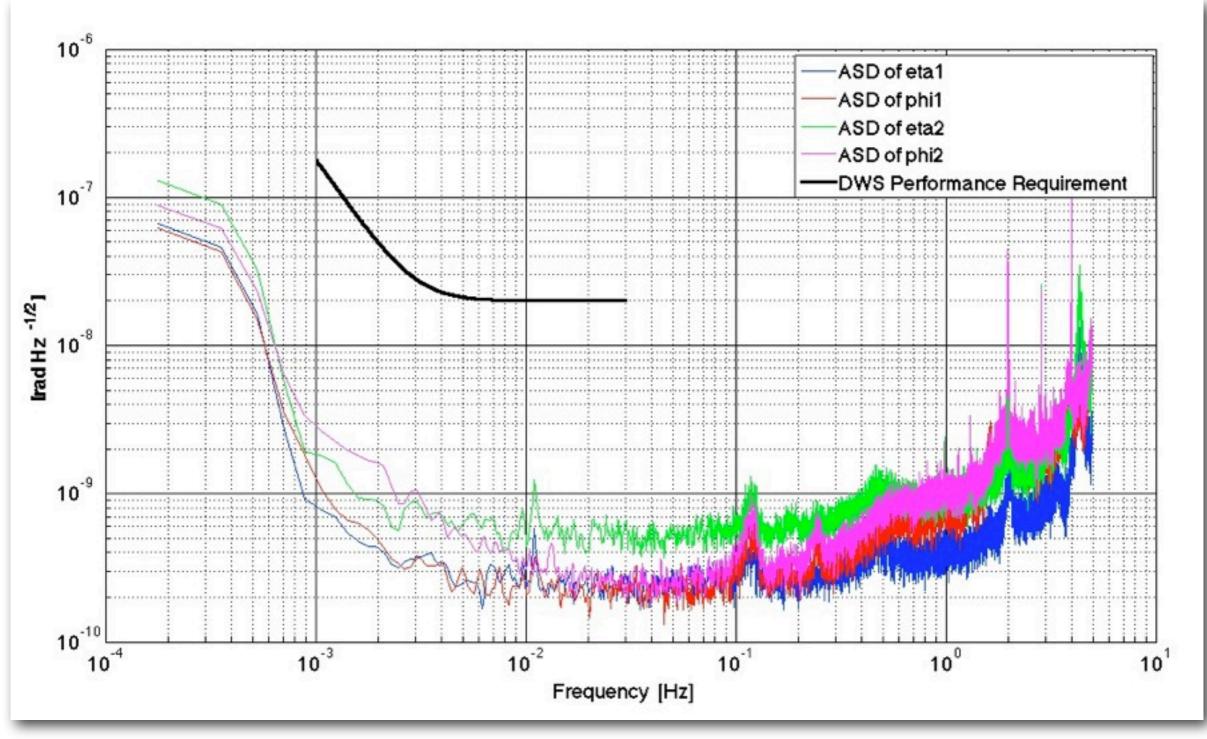




Interferometer displacement noise

OSTT OMS performance

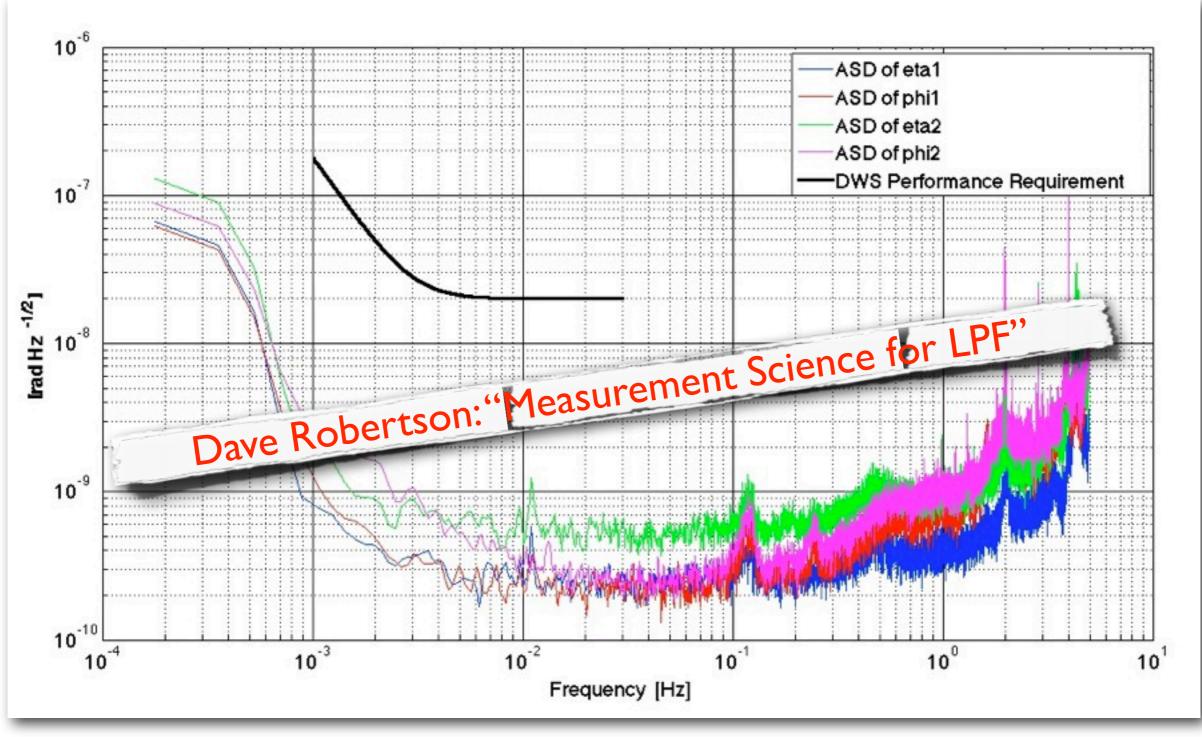




Interferometer angle sensing noise (Differential Wavefront Sensing)

OSTT OMS performance





Interferometer angle sensing noise (Differential Wavefront Sensing)

What's Missing? [1]



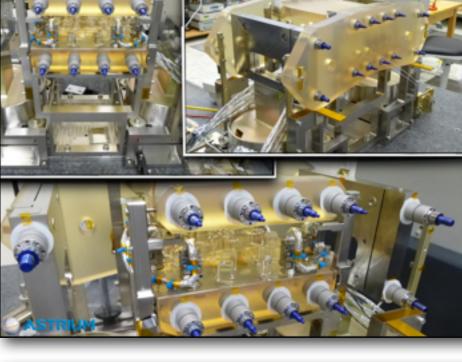
LTP Core Assembly

- All subsystems/components have been delivered
- The optical bench has been integrated to its support structure, and is now in storage awaiting delivery of the inertial sensors

Solution Ongoing Activities:

- Integration of the inertial sensors
- Following integration, the inertial sensors will be tested separately (thermal & vibration), prior to being integrated to the optical bench
- Final testing of the LCA will then take place followed by integration of the payload to the spacecraft







What's Missing? [1]



LTP Core Assembly

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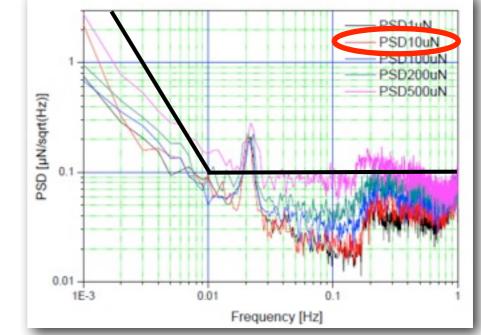
What's Missing? [2]

Microthrusters

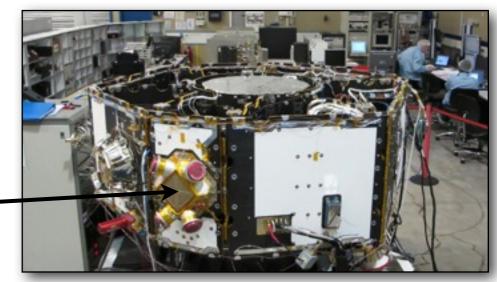
- After delays in the qualification schedule of the FEEP thrusters, we took the decision to change the thruster baseline to a Cold Gas micro-thruster
 - Recently developed for ESA's GAIA mission
- Performance of cold gas thrusters has improved several fold over the last few years
 - The GAIA cold gas thruster meets (and exceeds) the LPF requirements
 - Main issue is the low specific impulse of gas thrusters
 - Requires more kg of fuel per unit thrust

Colloid thrusters

- In addition to the Cold Gas Thrusters, LPF will also carry a set of Colloidal thrusters, as part of the NASA ST7 mission
- Colloids have been delivered and are integrated to the flight spacecraft



Measured thrust noise of GAIA thrusters





Microthrusters

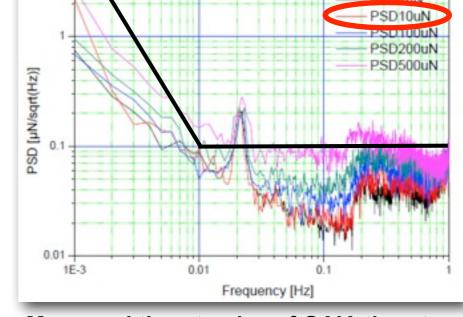
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Solloid thrusters

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Integrated to the flight spacecraft

Measured thrust noise of GAIA thrusters





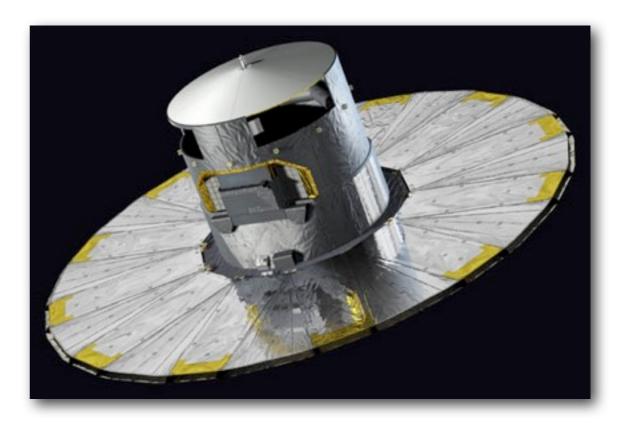




Cold Gas Thrusters

- Most flight hardware has been delivered to the industrial prime, and has gone through unit testing (e.g. magnetic testing)
- Final thruster head will be delivered at the end of this year
- Sold gas thrusters are now flying on GAIA
 - First performance results expected at the end of commissioning

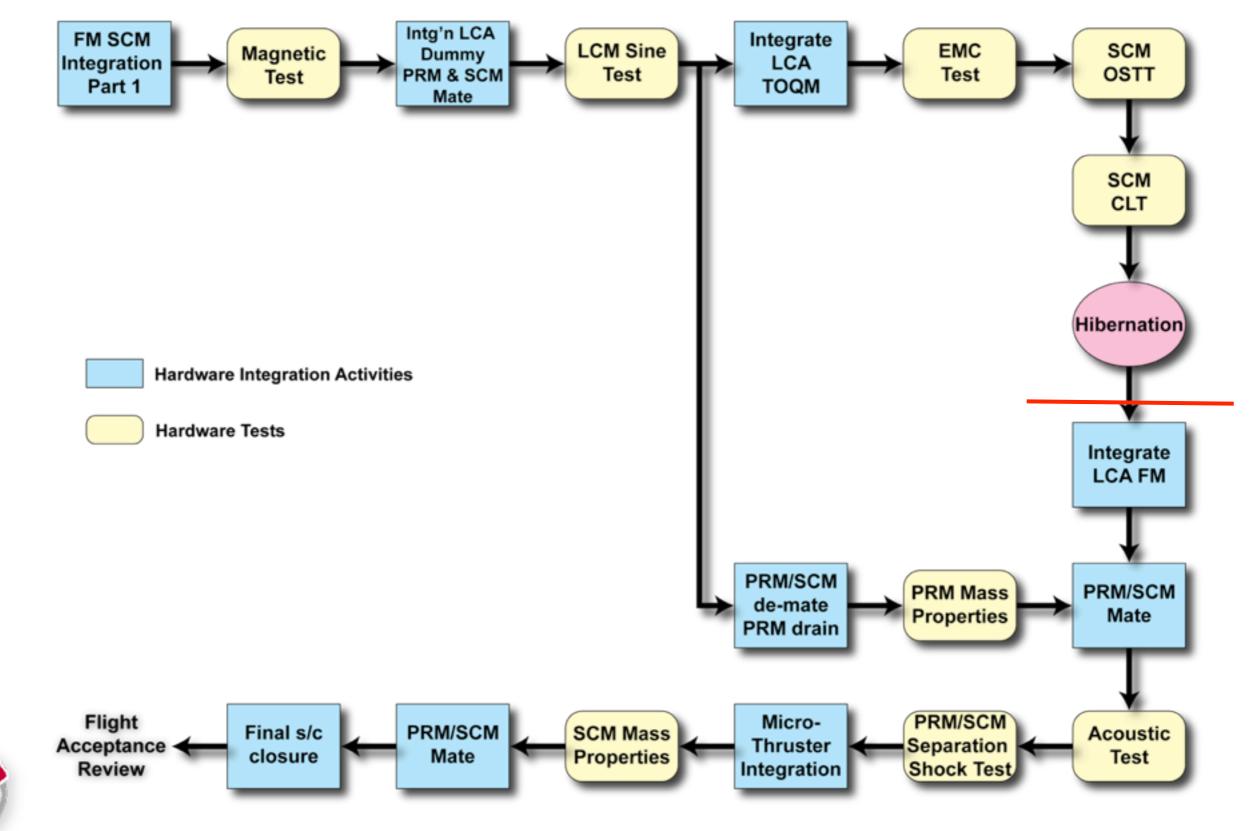






What's left? [1]





What's left? [2]



- In addition to the satellite system tests, several end-to-end tests will also be performed
- These tests will be performed on various platforms:
 - pre-SOVT test
 - These are a series of tests performed on a mission simulator
 - Tests are designed to prepare Payload Operation Requests (PORs) which can be used in subsequent tests (SOVT) and during operations
 - SOVT (System Operations Validation Test)
 - A test performed on the real-time testbed (RTB)
 - RTB is a *flat-sat* using the Engineering Models (EMs) of the hardware
 - The SOVT tests the end-to-end ground segment





What's left? [3]



End-to-end closed loop test (e2e CLT)

- This is a functional/performance test on the spacecraft
 - Special Check-Out Equipment is used where necessary
 - Inertial Sensors, Interferometer, Power System, Dynamics
- Designed to test the closedloop control of the s/c during science experiments



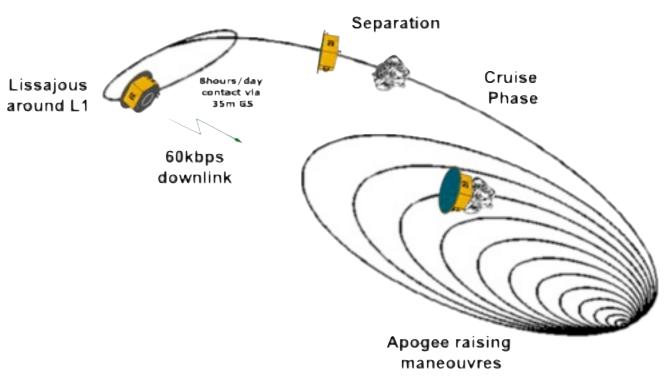


Launch and Operations



- LPF will be launched into a 200kmx1600km parking orbit by the new Arianespace launcher, VEGA
- 6 apogee raising manoeuvres required to deliver LPF to L1
- Prop module separates shortly before entering science orbit
- Final orbit is a ~500,000 x ~800,000km Lissajous orbit around L1







Operations



- Mission Operations are controlled from ESOC, Germany
- 8 hours ground station contact per day
 - 35m ESA Cebreros station
 - Downlink rate of 60kbps
- Science Operations controlled from ESAC, Spain
 - Operations run via Mission Timeline
- Real time commanding only during commissioning and contingency events















Conclusion



- LISA Pathfinder is well on its way to validate the technologies required by future spaceborne GW detectors
- Results from the OSTT demonstrate performance better than requirements
- The project is now out of hibernation, and preparing for the final environmental tests and the launch campaign
- Launch is scheduled for July 31st 2015
 - L 438 days and counting.....





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- Results from the OSTT demonstrate performance better than requirements
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- Launch is scheduled for July 31st 2015
 - L 438 days and counting.....
- The future is now....it's LISA Pathfinder





Thank you

ESA ESOC EADS Astrium ŪK EADS Astrium GmbH University of Trento Albert Einstein Institute University of Glasgow University of Birmingham Imperial College London ETH-Zurich

Institut d-Estudis Espacials de Catalunya Universidad Politecnica de Barcelona

Laben Carlo Gavazzi Space ALTA ARCS **Contrav** Kaiser Threde NTE SCISY Spacebe SRO echnolog ESA ARM

ASA Goddard

BUSEK

Monday 19 May 14

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