Recent developments in frequency domain multiplexing

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Outline

- Frequency domain multiplexed readout building blocks
- Building a prototype: European/Japanese EURECA collaboration
- Status building blocks
- Increasing the number of channels
- Summary



FREQUENCY-DOMAIN-MULTIPLEXING

Key characteristics:

- TESs act as AM-modulators of bias power
- Bias power sources separated in frequency space (i.e TESs AC-biased at frequencies *f1*, *f2*, *f3*,)
- LC band pass filter per pixel to prevent addition of wide-band noise
- Summed signal read-out by one SQUID-amplifier per column





FDM compatible TES bias circuit

Dissipation-less voltage source

Capacitive voltage divider, instead of classical resistive voltage division

- Functional equivalent with resistive divider
- > No extra dissipating element at base temperature
- Lower Q-requirements for capacitor





Building a prototype: the EURopEan Calorimeter Array (EURECA) Project

AIMS and Motivation:

- Design, build, and test of a prototype X-ray Imaging Spectrometer to demonstrate technical feasibility/readiness for a cryogenic space instrument by end 2007
- Use EURECA as a vehicle to breed a European/Japanese collaboration that is able and interested to deliver a cryogenic instrument for imaging X-ray spectroscopy
- Open up the potential to participate in future missions, like ESA's XEUS (>2020), NASA's Con-X (>2020), future Japanese missions like NEXT (2015) and DIOS (2012), Italian's Estremo (2015), Dutch NEW (2015)
- Acquire development funding at (multi) national level



EURECA Project

ADR Cooler

- Commercial ADR (Janis)
- Flight type ADR

Detectors

- Si-micromaching
- Development + tests
- Mo-based bilayers

LC-filters

• Alternative routes

SQUIDs

Three routes

Electronics

- LNA
- AC-BIAS + C&C
- Cold FLL
- Data Acquisition (BESSY)
- Data analysis software
 - System
 - Algorithms

Contibutions/Partners

PSI (Zürich) MSSL (London) →SRON MESA (UTwente) TMU (Japan), INFN(Genua), INAF(Rome), KIP (Heidelberg) IMM(Madrid), ICMA(Barcelona, Zaragossa) →SRON INA + ICMA (Zaragossa)

PTB (Berlin), VTT (Helsinki), SII (Tokyo) →SRON VTT (Helsinki) PSI (Zürich)

Alcatel Alenia Space (Milano) X-ray Astronomy (Leicester)

IFCA(Santander), MSSL(London), Astr. Obs (Geneva) X-ray Astronomy (Leicester)

Prototype development items

- EMI Shielding
- Microcalorimeters (tomorrow)
- LC bandpass filters
- SQUIDs
- Cryogenic detector unit
- SQUID front-end electronics
- AC bias source electronics
- Demultiplexing electronics
- Data analysis/processing

Approach:

- Conservative baseline system
- > lumped elements, which can be exchanged once better parts are available



EMI prevention

- EMI origin and sensitivity:
 - pickup by wiring
 - > 1 fW of power (broad band!) eq. ΔE of 1 eV
- Prevention approach:

Pickup attenuated by shielding and filtering in multiple stages

> Minimize complexity within shielded environment

- Implementation:
 - Filters at entrance FEE-box and at 4 K interconnections

Differential electronics and twisted wire-pairs to reject common mode disturbances

- > Separation between SQUID FLL electronics and other parts
- > PC's + external equipment coupled by optical links



Prototype topological overview



Cross section cold head shielding



Summing Topology

Cold Head Layout



Current Summing BCC at input 8 channels/squid TES-ARRAY



Japanese Ch. Flux Summing 8-input SQUID BCC via FB



LC-filters



SQUIDs

Cold head and COOLER



Cold head integration in progress





LC-filters



Filter Coil 0.75 x 0.75 mm²

LC-wafer for EURECA:

 $> C = 4.3 \text{ nF/mm}^2$

➤ Q (expected) = 10.000

Eureca design values

- $f = 0.8 2 MHz; \Delta f = 100 200 kHz$
- $L = 400 \text{ nH for } R = 40 \text{ m}\Omega$ •
- So C = 16 98 nF
- $Q > 800 \text{ f (MHz) for } R_{FSR} < 3.5 \text{ m}\Omega$

Dielectric material

- Nb and Ta-oxides show too much absorption
- > 25 nm Al₂O₃ with $\varepsilon_R = 9.8$ has been chosen



SQUID current amplifiers



VTT input SQUID $Ø_N = 0.12 \ \mu Ø_0 / \sqrt{Hz} @ 4K$ $L_{in} \approx 1nH$ $I_n = 3.5 \ pA / \sqrt{Hz}$ $T_N = 8 - 12 \ K (2^{nd} \ SQUID-array required)$





Artist impression of EURECA electronics





Functional diagram electronics

Applied technologies:

- > AC-bias generation + Bias current cancellation (BCC): DDS chips
- > DEMUX by ADC + digital processing: FPGA (later ASIC)
- Signal processing (photon energy extraction): FPGA + DSP





Functional diagram AC-bias sources

- 1 bias source/pixel (so f adjustable -relax of LCfilter f)
- +/-100 Hz < fc < +/-2 kHz
 <-120 dBc amplitude noise
 <-100 dBc phase noise
 < -80 dB spurs
- Digital (DDS-chips) meet the specifications
- FPGA + DAC's could as well
- ASICs under development for XEUS (power optimization)





Functional 8-channel AC-bias card



SRON

Functional diagram DEMUX-cards and C&C

- > 2 16-bit ADC's(25 MSA/s)
- Digital DEMUX in an FPGA
- Fast DAC's allow generation of signals for base-band feedback
- C&C card has large DSP to run energy extraction algorithm on data





EURECA Planning Overview



Qualification of a DC-biased pixel in dry ADR at BESSY September 2006

Start Integration 5 x 5 array + FDM-readout Autumn 2006

Initial testing (one channel) Begin 2007

Synchrotron testing (all channels) End 2007



Increasing the number of channels

Channels per SQUID

- Number of channels limited by bandwidth
- Bandwidth limited by FLL system
- Two methods to increase bandwidth
 - 1. Baseband feedback
 - 2. Coarse/fine amplifier topology

Channels per cm² electronics => Electronics miniaturization by ASICs

Remember: EURECA will have 8 channels/ SQUID, 24 channels in total



Dynamic range and Flux locked loop (FLL)

Requirements:

- 1. Signal dynamic range: $+/-5 \ 10^{6} \sqrt{Hz}$
- 2. Signal bandwidth: 150 200 kHz/pixel (crosstalk limited)

SQUID Dyn. Range:

- > Typical SQUID: +/-0.25 $\Phi_0/\Phi_N = 210^6 \sqrt{Hz}$
- > SQUID in FLL:

Assuming $G_{FLL} = 4$ (Unity-GBW, propagation delay limited \approx 10 MHz) => 2.5 MHz available, I.e. 9 - 12 channels/SQUID. This will be build for EURECA.

So for larger arrays, other systems than standard FLL required.



Bandwidth improvement by BASEBAND FEEDBACK

Carrier contains no information => feedback only on envelope

- A bandwidth up to about 10 MHz possible (SQUID back-action noise and LC-filter Q-factor limited)
- Stability issues limit packing density of channels
- \succ FLL-gain of 5x with 32 pixels between 1 10 MHz appears feasible





Bandwidth improvement by Coarse/Fine Amplifier Topology

Coarse SQUID

Operating principle:

- 2 SQUIDs: coarse and fine
- F-SQUID measures error made by coarse SQUID
- Subtraction of coarse and fine signal yields original signal
- Coarse/fine topology enables 10 MHz bandwidth (backaction noise limited)
 => 40 - 80 channels/squid (100 µs microcalorimeters)

Requirements:

- Coarse SQUID amplification sufficient to match RTamplifier
- A*S and A*E > RT amplifier noise => 2 stage SQUID
- > Largest signal < 0.5 ϕ_0
- Flux noise F-SQUID negligible => minimum M-value





Summary and Conclusions

- EURECA well under way with Preliminary Design Review in Jan.
 2006. Start integration 1st channel in ADR by end-2006
- Integration of single TES-pixel with DC-electronics in dry ADR started with aim to perform BESSY-calibrations in 3rd week of September 2006
- FDM readout can be implemented without the bias resistors, and their dissipation
- FDM with standard FLL-electronics will only multiplex about 10 pixels per SQUID-channel with XEUS requirements (E_{max} =10 keV, $\Delta E = 2 \text{ eV}$, and $\tau = 100 \text{ }\mu\text{s}$)
- Coarse/Fine amplifier topology, Base-band feedback, or a combinations should offer appreciably better performance. It is planned to start working on this by 2007 in parallel to EURECA
- SQUIDs close to the requirements available. But further optimization still required
- ASIC developments for Space (power reduction) is starting



COLLABORATORS ON EURECA

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