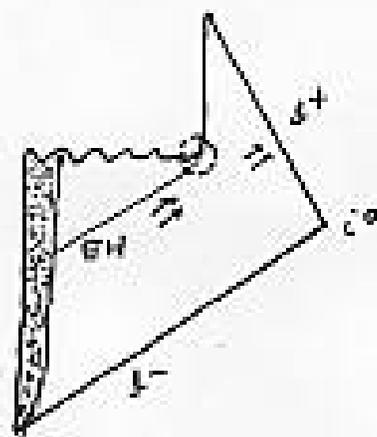


Quantum Geometry & BHs

Going beyond semi-classical approximation

Semi-classical scenario



QFT in CST
+
Heuristics of
back reaction

Limitations overlooked: cannot trust figure

- i) Near the entire singularity
- ii) Role of the EH: global, teleological notion

What really happens? Loop QG paradigm

1. Quantum geometry
2. Isolated horizons & Hawking process
3. Dynamical horizons
4. BH evaporation & Information loss

1. Quantum Geometry

- Gravity \sim geometry in GR ; No background metric. Matter & geometry, both "born" quantum mechanically.
- Quantum Riemannian geometry : New arena developed systematically by \sim 2 dozen groups in the mid-90s.

◦ gravity sector: $\underbrace{A_a^i}_{\text{SU(2) spin-conn.}}$; $\underbrace{E_i^a}_{\text{'Triad' / geometry}}$

$\mathcal{H} = L^2(\mathcal{A}, d\mu_0)$; \hat{h}_e, \hat{E}_i^a : well-defined

grav. holonomy.

\hat{A}_a^i : NOT defined
Non-Fock rep.

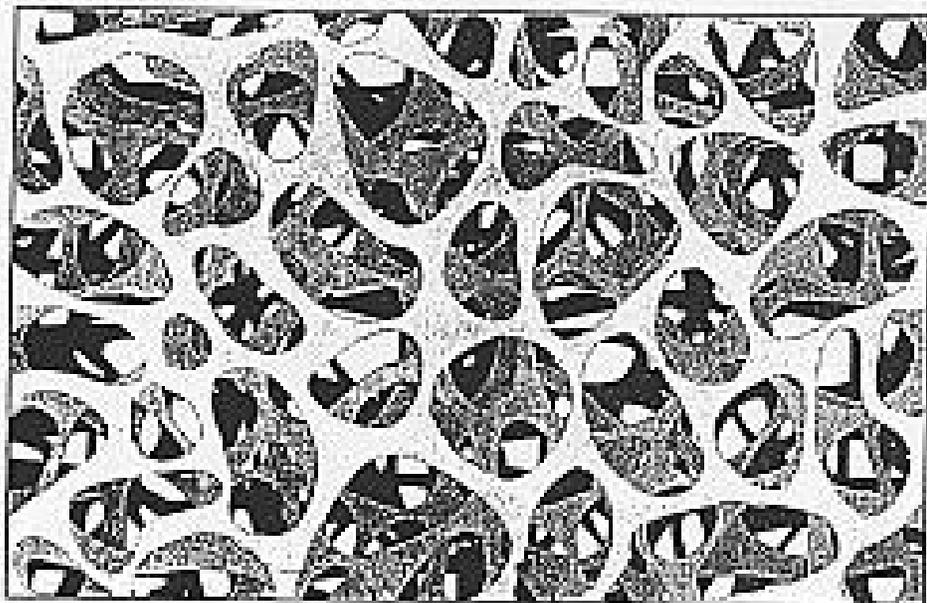
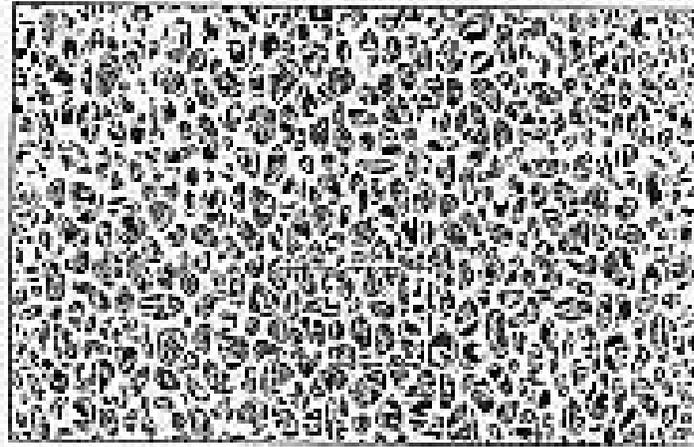
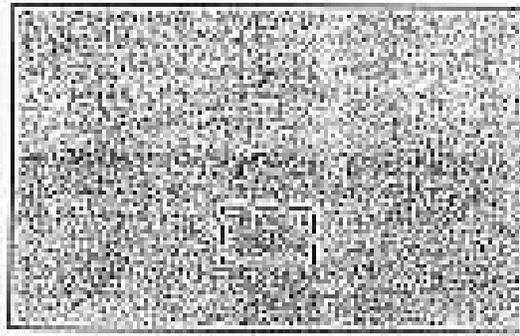
- Fundamental excitations : 1-dimensional 'polymer' geometry
"Flux lines of Area"



Technically : holonomies

$$a_s = \otimes \sqrt{S(S+1)} \ell_p^2$$

- Eigenvalues of all geometric operators : discrete
Analogy to H-atom : geometry quantized
Continuum only an approximation



1 Figure 29. Spacetime is smooth (top) at macroscopic scales but may be foamy at the Planck
 2 scale (middle and bottom). Quantum-mechanical fluctuations of the vacuum perpetually rip
 3 and join the fabric of spacetime, constantly changing its topology. These topological
 4 fluctuations constitute spacetime foam.

5 other (unknowable) universe. Conse-
 6 quently, a fundamental theory of ev-
 7 erything (such as superstrings) would
 8 never be able to predict the values of
 9 the coupling constants because worm-
 0 holes would wash these values away.

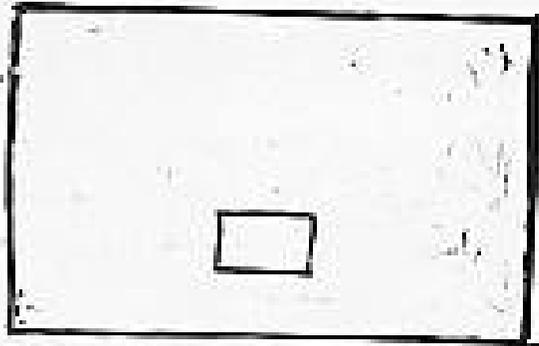
Interestingly, wormholes might pro-

vide a different mechanism that acts to
 fix the values of the coupling constants,
 including the cosmological constant.
 Coleman shows that the sum-over-his-
 tories corresponding to the state vector
 of the universe is dominated by a single
 term. The term corresponds to a uni-

Wheeler's
 qualitative
 picture:

Analogy with
 an ocean

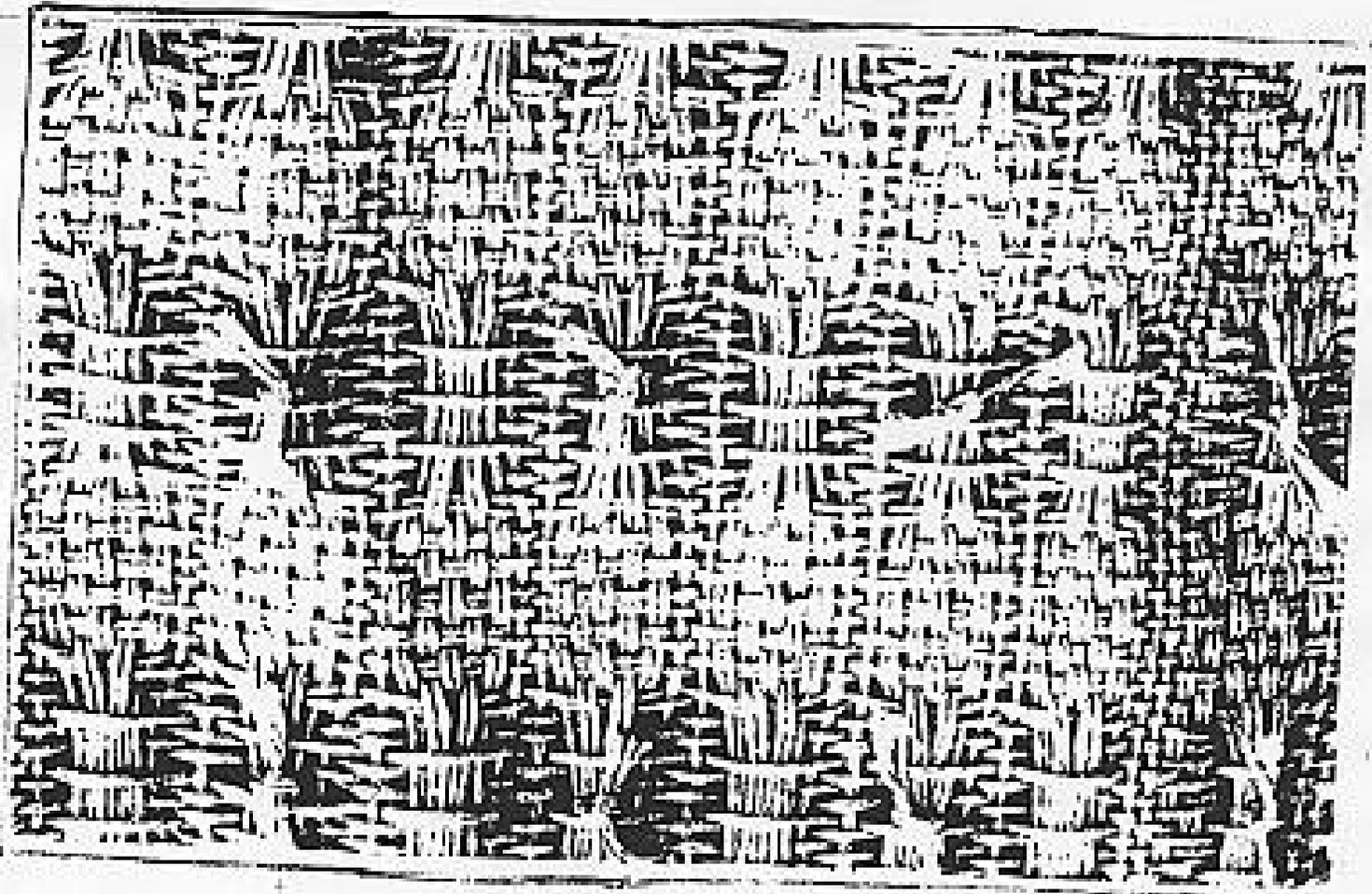
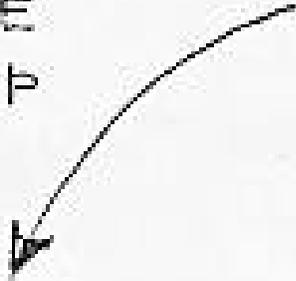
"space-time
 Foam"



Continuum

HUGE
Blow-up

$\sim 10^{34}$



Planck-scale
structure

List 1

The first 100 area eigenvalues of the quantum area operator are listed below. These numbers need to be multiplied by the planck length squared to give the actual areas.

1	0.43301269	51	1.63895792
2	0.50000000	52	1.70710677
3	0.56143781	53	1.71391368
4	0.70710677	54	1.72474492
5	0.82915622	55	1.73205078
6	0.86602539	56	1.75116259
7	0.93301269	57	1.75588831
8	0.96024586	58	1.76216891
9	1.00000000	59	1.77947181
10	1.08972478	60	1.78535712
11	1.09445750	61	1.79683155
12	1.11003401	62	1.79740208
13	1.14011946	63	1.79903005
14	1.16143781	64	1.80155727
15	1.19895792	65	1.80277562
16	1.20710677	66	1.82287562
17	1.22474492	67	1.82495376
18	1.26216891	68	1.82514077
19	1.29903805	69	1.82915622
20	1.32207562	70	1.83427125
21	1.32915622	71	1.84722623
22	1.36602539	72	1.85404956
23	1.36854458	73	1.86039573
24	1.39194107	74	1.86602539
25	1.40129856	75	1.86854458
26	1.41421354	76	1.87002075
27	1.43101269	77	1.88618273
28	1.46824586	78	1.89194107
29	1.47902000	79	1.90125856
30	1.49059113	80	1.90606469
31	1.50000000	81	1.91203269
32	1.52273747	82	1.91421354
33	1.52746320	83	1.91800100
34	1.53626299	84	1.92020642
35	1.55104670	85	1.92160672
36	1.56124949	86	1.93105169
37	1.57313216	87	1.93301269
38	1.58113885	88	1.93649173
39	1.58972478	89	1.94719023
40	1.59445050	90	1.95575017
41	1.61803401	91	1.96047586
42	1.62968367	92	1.96824586
43	1.63197061	93	1.96927560
44	1.63935939	94	1.97902000
45	1.64011946	95	1.98405939
46	1.65775761	96	1.98431349
47	1.65031244	97	1.99059403
48	1.66143781	98	1.99426219
49	1.67535792	99	2.00000000
50	1.69518751		

(Steve Fairhurst)

Eigenvalues
"crowd"

Geometry is
quantized and
in a very subtle
way.

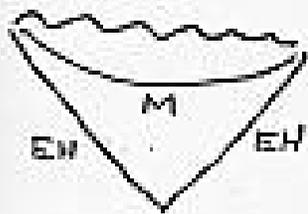
Spectra of
all geometric
Operators:
discrete!

(eigenvalues of \hat{A}_S)

$$a_s = 2\pi r \ell_p^2 \sum_i [2J_i^d (J_i^d + 1) + 2J_i^u (J_i^u + 1) - J_i^{u+d} (J_i^{u+d} + 1)]^{1/2}$$

$\mu \quad |d| : \frac{1}{2}$ integers; $J_i^{u+d} \in \{J_i^u + J_i^d, \dots, |J_i^u - J_i^d|\}$

Resolution of BH singularity



Schwartzschild interior-
(generic, 'stable' singularity)

- $M = \mathbb{S}^2 \times \mathbb{R}$; 4 kvFs
KS Minisuperspace.

2 DOF:

τ : Radius of \mathbb{S}^2
 ϕ : anisotropy.

Apply methods from QC
Mimic full, unreduced theory

New quantum mechanics!

$$\mathcal{H} \neq L^2(\mathbb{R}^2, d^2x) ; \quad \mathcal{H} = L^2(\bar{\mathbb{R}}_{\text{holo}}^2, d\mu_0)$$

only holonomies well-defined, not connections
[In QM: $\widehat{\exp i\alpha x}$ - fine but NOT \hat{x} !]

Inequivalence has dramatic consequences

co-triad diverges classically at the singularity.
Quantum co-triad operator bounded above!
 \Rightarrow curv bounded above

Quantum evolution: "WDW" equation

$$|\Psi\rangle = \sum_{\tau} \Psi_{\tau}(\phi) |\tau, \phi\rangle$$

$$= C^+(\tau) \hat{O}_+ \Psi_{\tau+4\tau_0}(\phi) + C^0(\tau) \hat{O}_0 \Psi_{\tau}(\phi) + C^-(\tau) \hat{O}_- \Psi_{\tau-4\tau_0}(\phi)$$

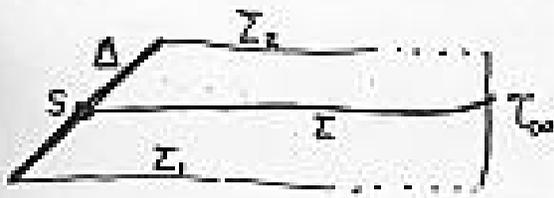
'Evolution' through discrete time steps.

Can evolve right through $\tau=0$!

classical space-time!

2. Quasi-local substitutes for event horizon

Time independent phase : Isolated Horizon



Δ : Null, $S^2 \times \mathbb{R}$, Intrinsic geometry time independent

\Rightarrow cross-sections S are marginally trapped S^2

IH Boundary Conditions \Rightarrow

- Well-defined action principle
- Hamiltonian framework
- Time evolution Hamiltonian \Leftrightarrow

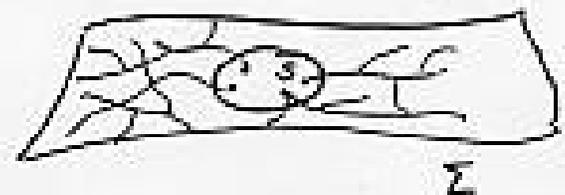
$$\delta E_o^{(t)} = \frac{k_a^{(t)}}{8\pi G} \delta Q_a + \int \Omega_a^{(t)} \delta J_a + \int \Phi_a^{(t)} \delta Q_a \quad \text{1st law}$$

Entropy \leftrightarrow Area

(AA, Beetle, Fairhurst, Krishnan, Lewandowski)

Quantum geometry of IHs

(AA, Baez, Krasnov)

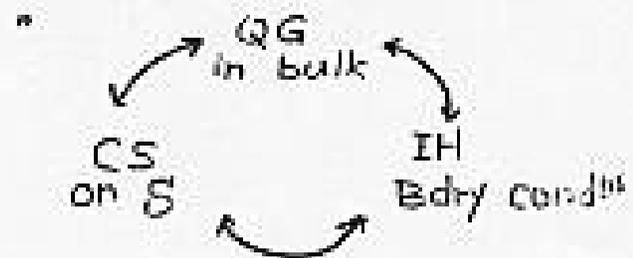


• strategy

• Bulk : Polymer geometry

• Horizon : U(1) Chern-Simons theory on punctured S^2

• Quantum-horizon Boundary condition:
very strong constraint

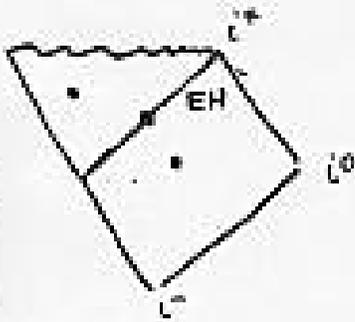


• Surface states account for entropy of all IHs

Distorted, rotating, colored, - non-minimal couplings,

\Downarrow
Coherent theory

• Dynamical horizons

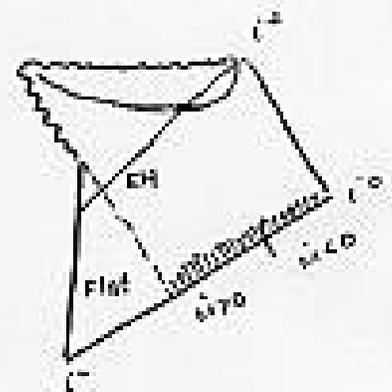
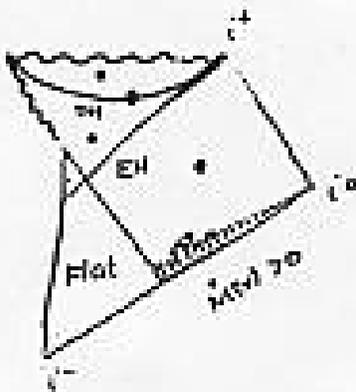


External black hole

- untrapped
- marginally trapped
- Trapped

Dynamical situations : A DH is $\mathbb{S}^2 \times \mathbb{R}$, foliated by MT_s ($\Theta_{(t)} = 0$, $\Theta_{(m)} < 0$).

Ex: Vaidya Solutions (spherical, null fluid)



DH: space-like

Time-like

Area of MT_s Increases

Decreases

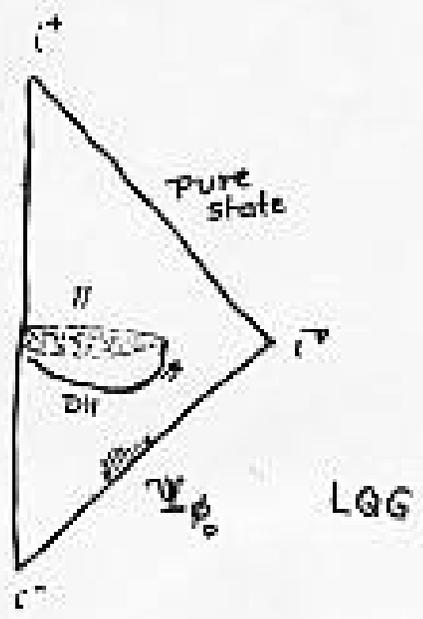
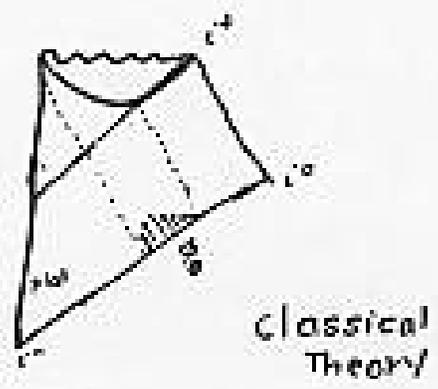
Energy flux +ve

-ve

Integral generalization of first law holds on general DH_s

4. BH evaporation: Quantum geometry paradigm

- Gather these precise results to develop a qualitative heuristic picture within LQG.
- Spherical collapse of a scalar field: Large BH



Deep Planck regime
No classical space-time

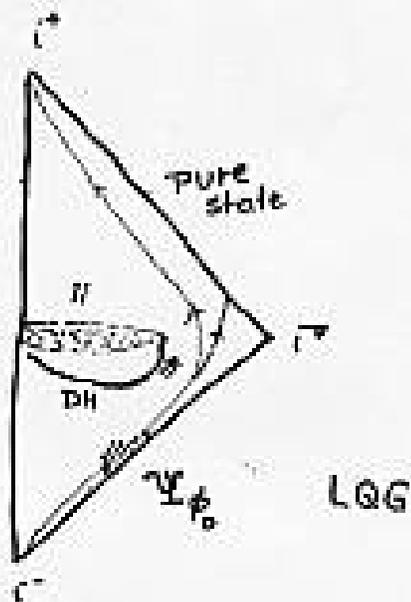
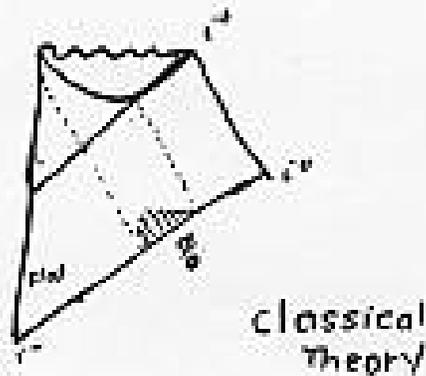
Assumptions:

i) Full quantum state becomes semi-classical again

ii) QNH = QTH when area $\sim \ell^2$

BH evaporation: Quantum geometry paradigm

- Gather these precise results to develop a qualitative - heuristic picture within LQG.
- Spherical collapse of a scalar field: Large BH



- Space-time does not end in a singularity \Rightarrow No sink of information.
- pure states on $f^- \mapsto$ pure states on f^+

• 'event horizon': no longer a useful notion

- DHs form in semi-classical region and evaporate.

$$\frac{dR}{dt} = -8\pi G R^2 \cdot T_{ab} l^a l^b$$

Area quanta \mapsto ϕ quanta

- Recovery of correlations (patience!)

⊗ Deep Planck regime
No classical space-time

Assumptions:

i) Full quantum state becomes semi-classical again

ii) ODH \rightarrow OTH when $area \gg l^2$

Summary.

Quantum geometry, underlying loop quantum gravity introduces qualitatively new elements in BH physics:

- Quantum isolated horizons described by U(1) CS theory on punctured sphere
- BH entropy for physical BHs (distortion, rotation, hair, non-minimal couplings, ...)
- BH singularity resolved.

When combined with classical IH and DH frameworks, one obtains a new paradigm for BH evaporation:

- No information loss
- Evaporation of DHs \rightsquigarrow Hawking radiation
- Elements of semi-classical scenario ok in suitable approximations
- Lorentzian picture of the process.