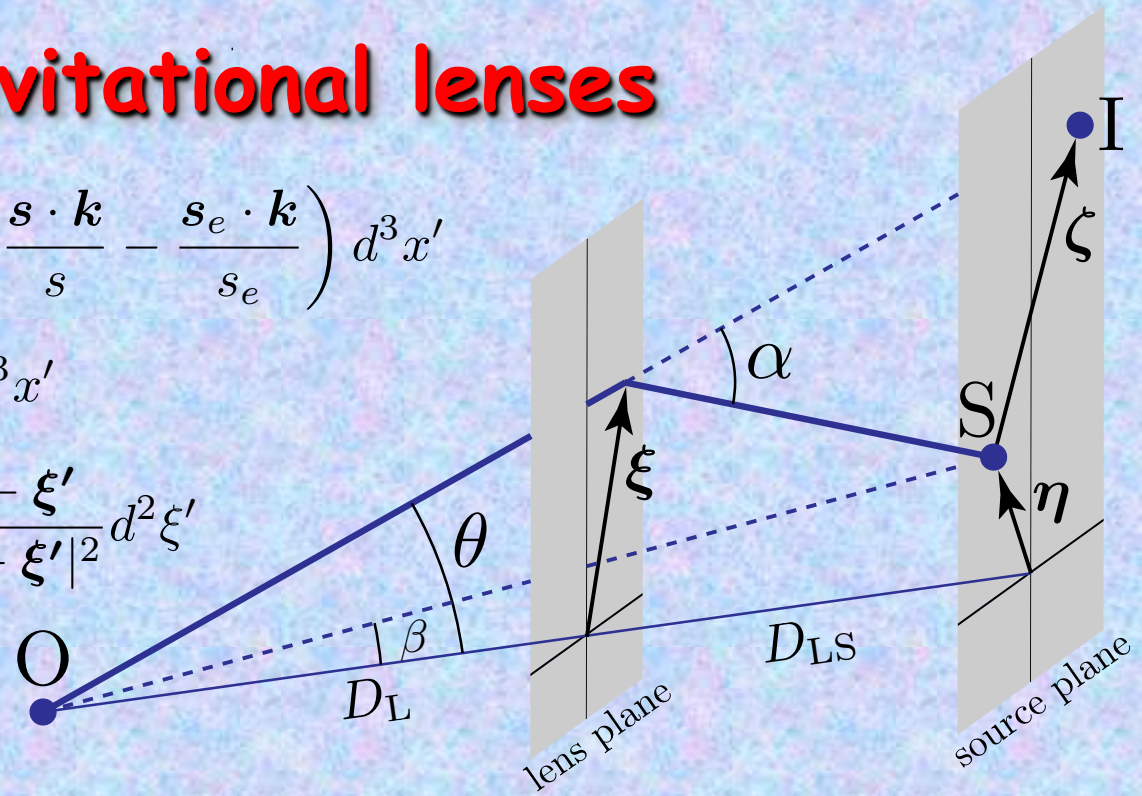


Gravitational lenses

$$\begin{aligned}\alpha(t) &= -\frac{2G}{c^2} \int \rho(t, \mathbf{x}') \frac{\mathbf{b}}{b^2} \left(\frac{\mathbf{s} \cdot \mathbf{k}}{s} - \frac{\mathbf{s}_e \cdot \mathbf{k}}{s_e} \right) d^3 x' \\ &= -\frac{4G}{c^2} \int \rho(t, \mathbf{x}') \frac{\mathbf{b}}{b^2} d^3 x' \\ &= -\frac{4G}{c^2} \int \Sigma(t, \boldsymbol{\xi}') \frac{\boldsymbol{\xi} - \boldsymbol{\xi}'}{|\boldsymbol{\xi} - \boldsymbol{\xi}'|^2} d^2 \xi'\end{aligned}$$



Lens equation

$$\theta + \frac{D_{LS}}{D_S} \alpha = \beta$$

$$\xi = D_L \theta$$

Schwarzschild lens

$$\alpha(\xi) = -\frac{4GM}{c^2} \frac{\xi}{\xi^2}$$

$$D_S \theta = \eta + \zeta = D_S \beta - D_{LS} \alpha$$

$$\theta - \frac{\theta_E^2}{\theta} = \beta$$

$$\theta_E^2 := \frac{4GM}{c^2} \frac{D_{LS}}{D_S D_L}$$



Gravitational lenses

$$\theta_{\pm} = \frac{1}{2} \left(\beta \pm \sqrt{\beta^2 + 4\theta_E^2} \right)$$

$$\theta_E \sim 2 \text{ as} \quad M \sim 10^{12} M_{\odot}$$

$$\sim 0.5 \text{ mas} \quad M \sim M_{\odot}$$

$$\beta \ll \theta_E, \quad \theta_{\pm} = \pm\theta_E + \beta/2$$

$$\beta = 0, \quad \theta_{\pm} = \pm\theta_E \quad \text{Einstein ring}$$

$$\beta \gg \theta_E, \quad \theta_{\pm} = \begin{cases} \beta + \theta_E^2/\beta \\ -\theta_E^2/\beta \end{cases}$$



Gravitational lenses: Magnification

$$\theta_{\pm} = \frac{1}{2} \left(\beta \pm \sqrt{\beta^2 + 4\theta_E^2} \right)$$

Brightness proportional to $\int \beta d\beta d\varphi$ at the source, $\int \theta d\theta d\varphi$ at the observer

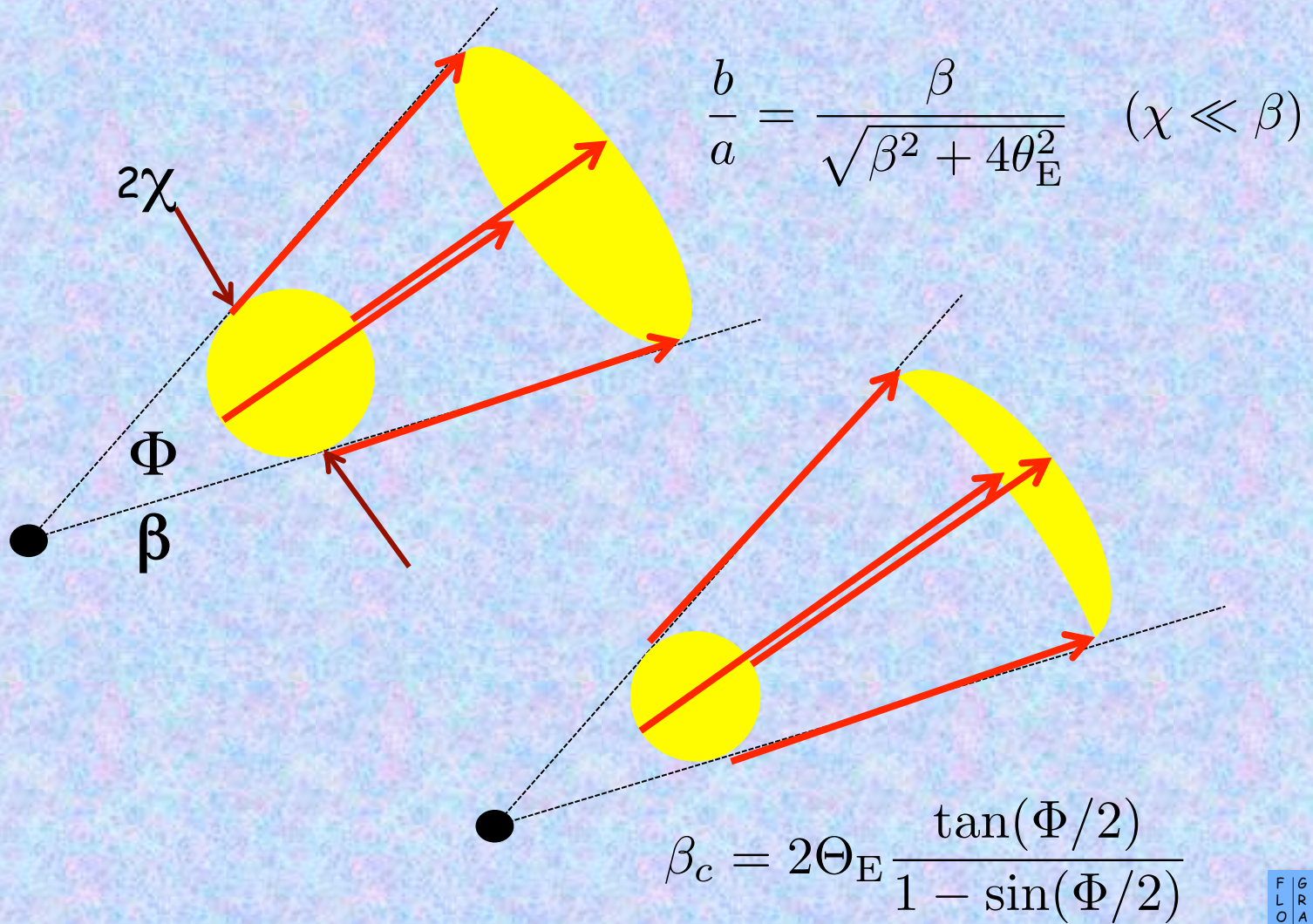
$$\mu_{\pm} = \frac{\theta_{\pm} d\theta_{\pm}}{\beta d\beta} = \pm \frac{1}{4} \left(\frac{\beta}{\sqrt{\beta^2 + 4\theta_E^2}} + \frac{\sqrt{\beta^2 + 4\theta_E^2}}{\beta} \pm 2 \right).$$

Microensing

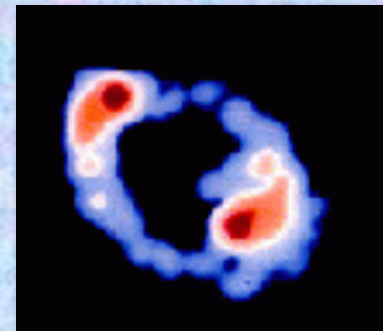
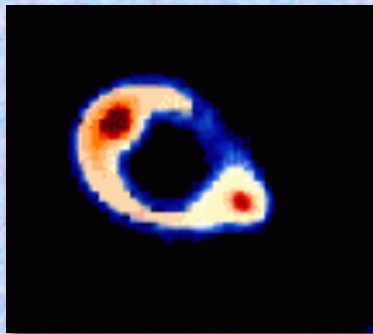
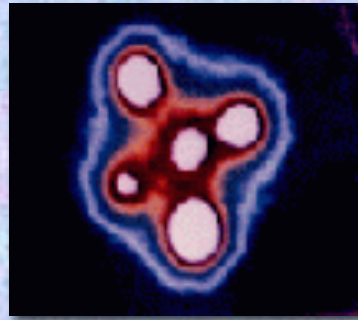
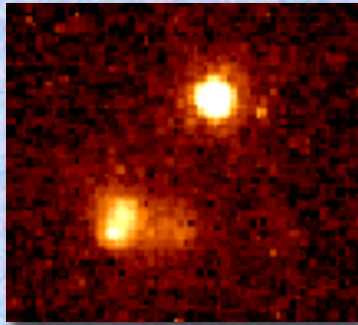
$$|\mu_+| + |\mu_-| = \frac{1}{2} \left(\frac{\beta}{\sqrt{\beta^2 + 4\theta_E^2}} + \frac{\sqrt{\beta^2 + 4\theta_E^2}}{\beta} \right) > 1$$



Gravitational lenses: Distortion



Gravitational lenses



Post-Newtonian time

Clock synchronization on the Earth

$$d\tau = \left[1 - \frac{1}{c^2} \left(\frac{1}{2} \bar{v}^2 + U \right) + O(c^{-4}) \right] d\bar{t}$$

$$= \left[1 - \frac{1}{c^2} \left(\frac{1}{2} v^2 + \mathbf{v} \cdot (\boldsymbol{\omega} \times \mathbf{r}) + \Phi \right) + O(c^{-4}) \right] d\bar{t}$$

$$\Phi = U + \frac{1}{2} \omega^2 (x^2 + y^2) \quad \text{Geoid potential}$$

Two clocks at rest

$$\tau_A = \left(1 - \frac{\Phi_{\text{geoid}}}{c^2} \right) \bar{t}$$

$$\tau_B = \left(1 - \frac{\Phi_{\text{geoid}}}{c^2} \right) \bar{t} + \tau_B^0$$

For travelling clock with $v \ll c$

$$\tau_{\text{trav}} = \tau_A(\bar{t}_1) + \left(1 - \frac{\Phi_{\text{geoid}}}{c^2} \right) (\bar{t}_2 - \bar{t}_1) - \frac{1}{c^2} \int_A^B (\boldsymbol{\omega} \times \mathbf{r}) \cdot d\mathbf{r}$$

$$\tau_B^0 = -\frac{1}{c^2} \int_A^B (\boldsymbol{\omega} \times \mathbf{r}) \cdot d\mathbf{r}$$

Sagnac effect:
207 ns for closed path
at the equator



Post-Newtonian time: GPS

$$d\tau_{\text{ground}} = \left(1 - \frac{\Phi_{\text{geoid}}}{c^2} \right) d\bar{t}$$

$$\begin{aligned} d\tau_{\text{sat}} &= \left[1 - \frac{1}{c^2} \left(\frac{1}{2} v^2 + U \right) \right] d\bar{t} \\ &= \left(1 - \frac{3 Gm}{2 c^2 a} \right) d\bar{t} \end{aligned}$$

$$d\tau_{\text{sat}} \left(1 + \frac{\Phi_{\text{geoid}}}{c^2} - \frac{3 Gm}{2 c^2 a} \right) d\tau_{\text{ground}}$$

$$\sim \frac{Gm}{c^2 a_{\oplus}}$$

$$\sim 6.9 \times 10^{-10}$$

$$\sim \frac{3 Gm}{2 c^2 (4.2 a_{\oplus})}$$

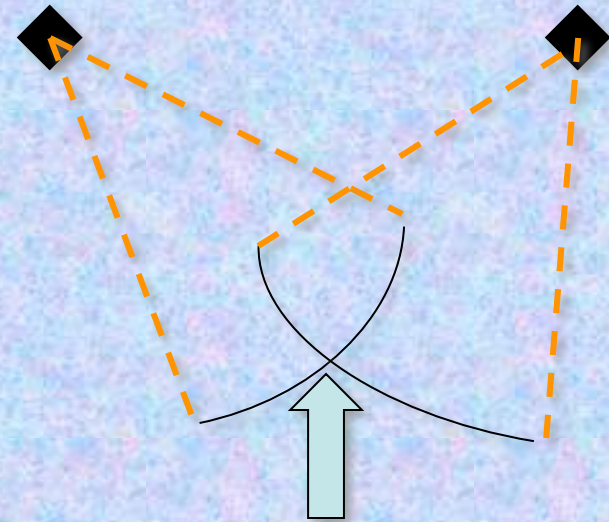
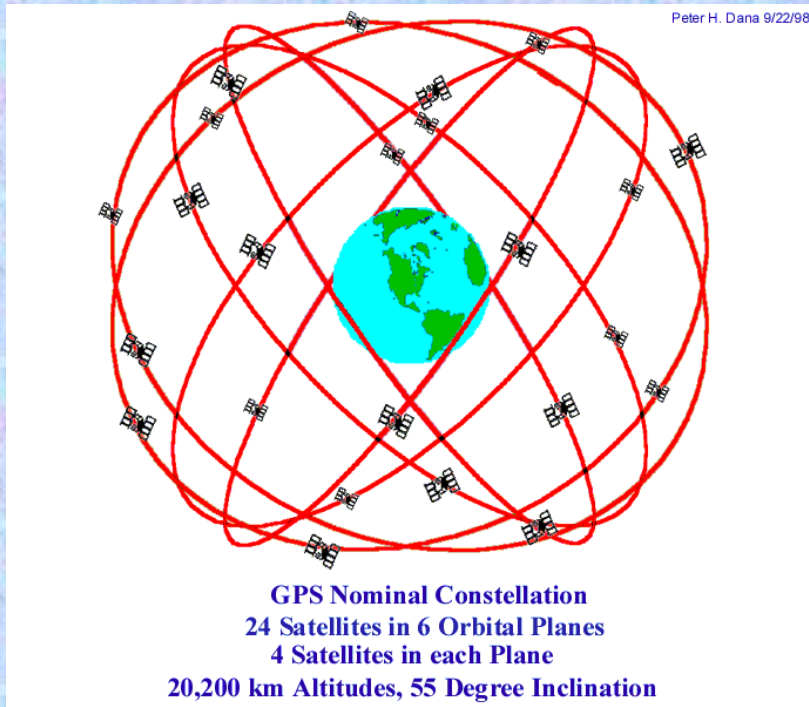
$$\sim 2.5 \times 10^{-10}$$

for GPS satellites



General Relativity and Daily Life

The Global Positioning System (GPS)



Navigation Requirement: 15 m \Rightarrow 50ns

Relativistic effects: 39,000 ns per day!

GR = 46,000

SR = -7,000

Relativity **must** be taken into account, for GPS to function