

$$\Lambda(v) = \Lambda(v_2) \Lambda(v_1) \quad (\text{two Lorentz boosts})$$

$$\begin{pmatrix} \gamma & -\beta\gamma \\ -\beta\gamma & \gamma \end{pmatrix} = \begin{pmatrix} \gamma_2 & -\beta_2\gamma_2 \\ -\beta_2\gamma_2 & \gamma_2 \end{pmatrix} \begin{pmatrix} \gamma_1 & -\beta_1\gamma_1 \\ -\beta_1\gamma_1 & \gamma_1 \end{pmatrix} = \gamma_1\gamma_2 \begin{pmatrix} 1+\beta_1\beta_2 & -\beta_1-\beta_2 \\ -\beta_1-\beta_2 & 1+\beta_1\beta_2 \end{pmatrix}$$

$$= \gamma_1\gamma_2 (1+\beta_1\beta_2) \begin{pmatrix} 1 & -\frac{\beta_1+\beta_2}{1+\beta_1\beta_2} \\ -\frac{\beta_1+\beta_2}{1+\beta_1\beta_2} & 1 \end{pmatrix}$$

read off:

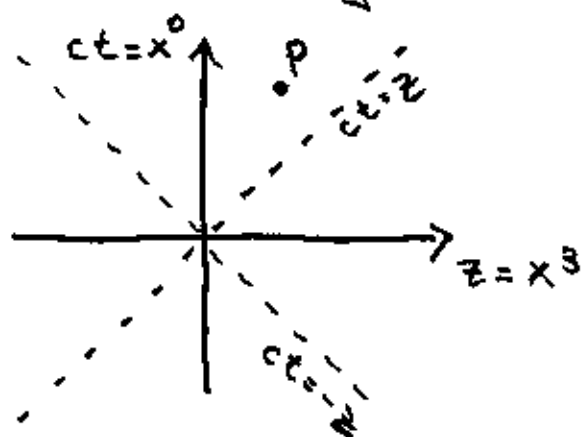
$$\beta = \frac{\beta_1 + \beta_2}{1 + \beta_1\beta_2}$$

$$v = \frac{v_1 + v_2}{1 + \frac{v_1 v_2}{c^2}} \quad (\text{relativistic addition of velocities})$$

Note: Lorentz boost leaves c invariant

$$\frac{v_1 + c}{1 + \frac{v_1 c}{c^2}} = \frac{v_1 + c}{1 + \frac{v_1}{c}} = c \quad \frac{v_1 + c}{v_1 + c} = c$$

Minkowski diagrams (spacetime diagrams)



special line:
light signal

$$c^2 t^2 - z^2 = 0$$

$$\pm z = ct$$

moving reference frame (Lorentz boost in z dir.)

Σ' at $t = t' = 0$, same origins $z = z' = 0$

coordinate axis of Σ' ,

a) $0 = z' = \gamma(z - vt)$

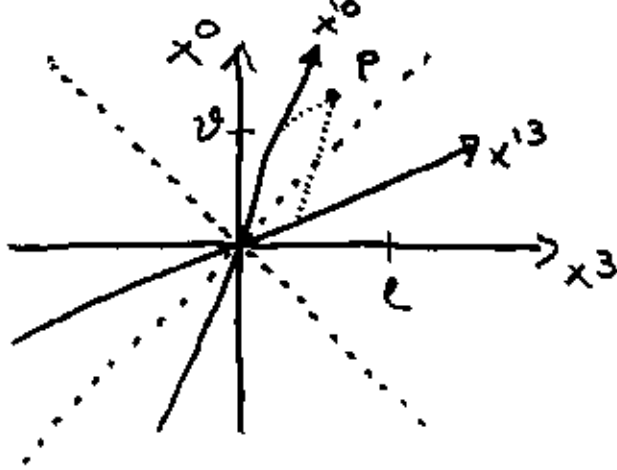
$\Rightarrow z = vt$

$\Leftrightarrow ct = \frac{c}{v}t = \frac{z}{\beta}$ (condition for ct' axis)

b) $0 = ct' = c\gamma(t - \frac{v}{c^2}z)$

$\Rightarrow t = \frac{v}{c^2}z$

$\Leftrightarrow ct = \frac{v}{c}z = \beta z$ (condition for z' axis)



scaling of axis using invariant (distance)

$$s^2 = x^\mu x_\mu = c^2 t^2 - (x^2 + y^2 + z^2)$$

for "scale" on x^3 axis, we have $s^2 = -l^2$

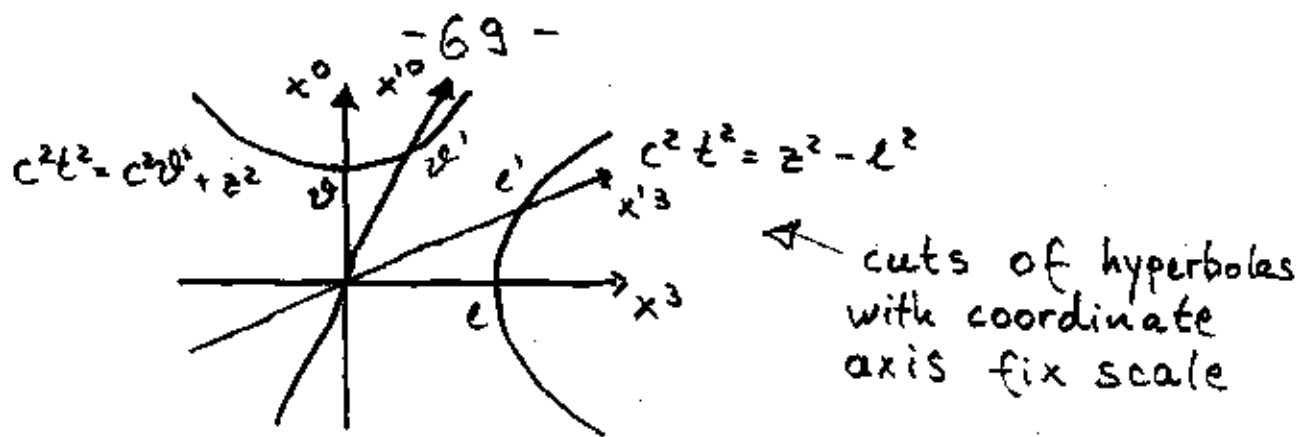
thus all points with $c^2 t^2 - z^2 = -l^2$

are invariant: $(ct)^2 = z^2 - l^2$ (hyperbola in Minkowski diagram)

for "scale" z' on x^0 axis, we have $s^2 = c^2 z'^2$

all points with $c^2 t^2 - z^2 = c^2 z'^2$

are invariant $c^2 t^2 = c^2 z'^2 + z^2$ (hyperbola)



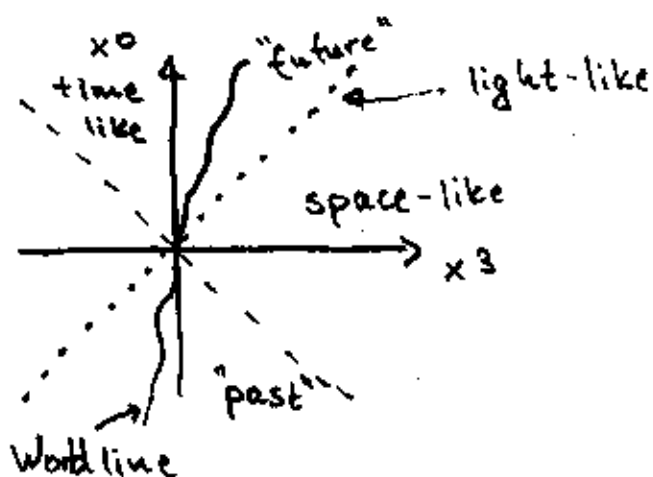
types of 4-vectors, worldline

$$s^2 = x^\mu x_\mu = c^2 t^2 - (x^2 + y^2 + z^2)$$

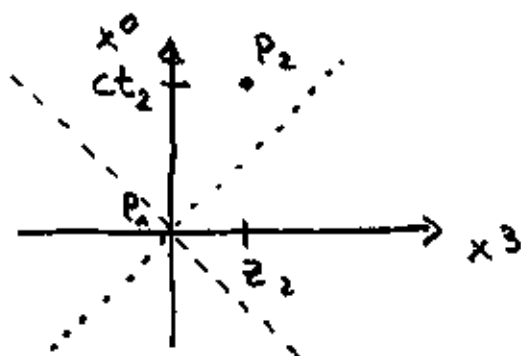
time-like 4 vector $s^2 > 0$, $|ct| > |\vec{r}|$

light-like 4 vector $s^2 = 0$, $|ct| = |\vec{r}|$

space-like 4 vector $s^2 < 0$, $|ct| < |\vec{r}|$

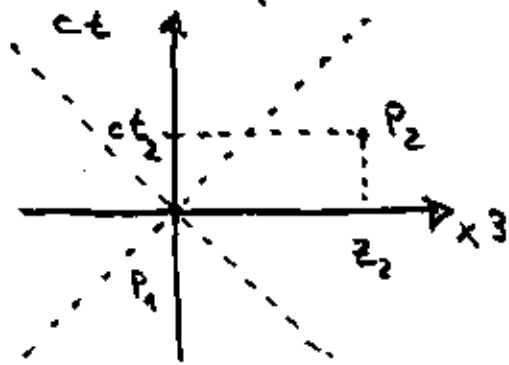


events with time-like distance $s^2 > 0$



exists reference frame in which P_1 and P_2 is at the same place:
 boost with $\beta = \frac{z_2}{ct_2}$

events with space-like distance

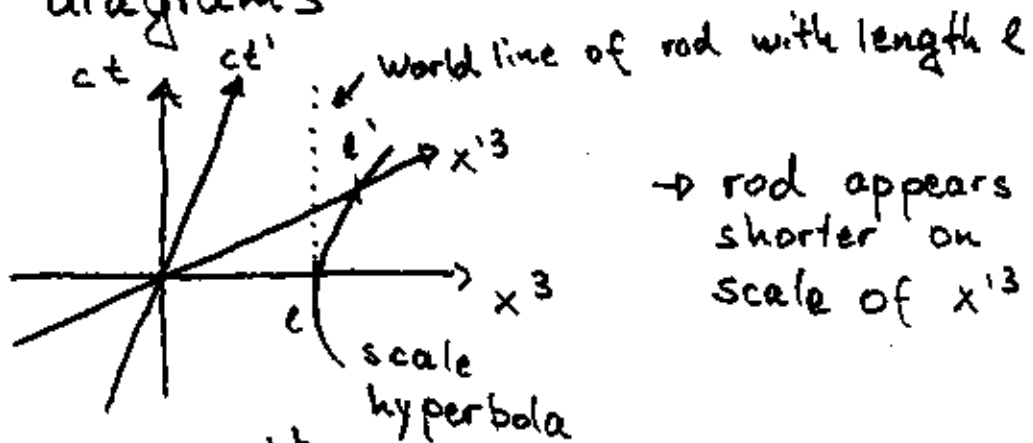


Lorentz boost with $\beta = \frac{ct_2}{z_2}$

(events at the same time)

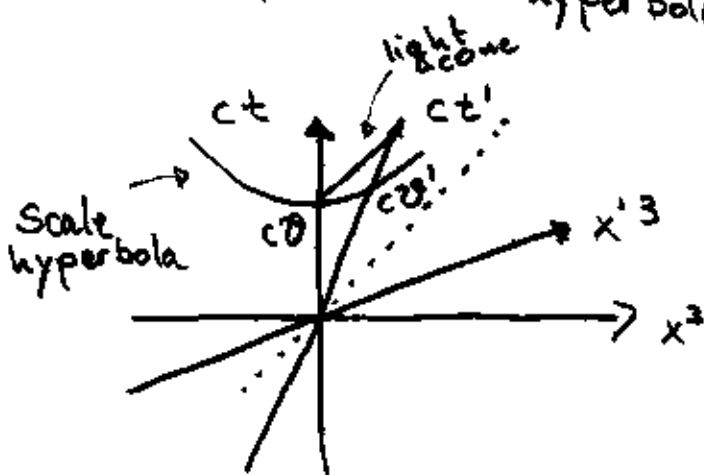
Length contraction and time dilatation in Minkowski diagrams

1)



→ rod appears shorter on scale of x'^3 axis

2)



light cone of event at $x^3 = 0$ reaches ct' axes at time $> ct_2'$

Covariant formulation of classical mechanics

goal: write physical laws such that these are invariant under Lorentz transform

note: $\vec{F} = \dot{\vec{p}}$ is only Galilei invariant

Reminder: tensors of n th rank

$n = 0$ scalars invariant

example: $\vec{r} \cdot \vec{r} = r^2$ in \mathbb{R}^n

$ds^2 = dx^\mu dx_\mu$ in Minkowski space

$n = 1$ vectors (tensors of 1st rank)

example $\vec{r} \rightarrow \vec{r}' = D \vec{r} = \sum_{i,j=1}^3 d^{ij} x^j \vec{e}_i$

$$x^\mu \rightarrow x'^\mu = \Lambda^\mu_\nu x^\nu \quad (*)$$

$n = 2$ tensors

example inertial tensor

$$J^{\ell m} = \int d^3 \vec{r} (\delta^{\ell m} \vec{r}^2 - x^\ell x^m) \rho(\vec{r})$$

$$J^{\ell m} \rightarrow J'^{\ell m} = (D J D^T)^{\ell m}$$

$$F^{\mu\nu} \rightarrow F'^{\mu\nu} = \Lambda^\mu_\alpha \Lambda^\nu_\beta F^{\alpha\beta}$$

Lorentz transformation as partial derivative:

$$\frac{\partial x'^\mu}{\partial x^\nu} = \Lambda^\mu_\nu \quad (\text{use } *)$$

alternatively: $(\Lambda^{-1})^\nu_\mu x'^\mu = \overbrace{(\Lambda^{-1})^\nu_\mu}^{\delta^\nu_{\nu'}}$ $\Lambda^\mu_{\nu'} x^{\nu'}$

$$\Rightarrow \frac{\partial x^\nu}{\partial x'^\mu} = (\Lambda^{-1})^\nu_\mu$$

contravariant vector

$$x'^{\mu} = \Lambda^{\mu}_{\nu} x^{\nu}$$

covariant vector

$$x'_{\mu} = (\Lambda^{-1})^{\nu}_{\mu} x_{\nu} = x_{\nu} (\Lambda^{-1})^{\nu}_{\mu} \quad (\text{transforms as coordinate axis})$$

check correctness by Lorentz invariance:

$$x'_{\mu} x'^{\mu} = x_{\nu} (\Lambda^{-1})^{\nu}_{\mu} \Lambda^{\mu}_{\nu'} x^{\nu'} = x_{\nu} g^{\nu}_{\nu'} x^{\nu'} = x_{\nu} x^{\nu}$$

tensors of 2nd rank

contravariant: $F^{\mu\nu}$

mixed: F^{μ}_{ν}

covariant: $F_{\mu\nu}$

each index transforms as contravariant vector

Tensor contraction:

$$F^{\mu}_{\nu} \rightarrow F^{\mu}_{\mu} = \sum_{\mu=0}^3 F^{\mu}_{\mu}$$

transformation: $F^{\mu}_{\mu} \rightarrow F'^{\mu}_{\mu} = \Lambda^{\mu}_{\alpha} F^{\alpha}_{\beta} (\Lambda^{-1})^{\beta}_{\mu}$

$$= (\Lambda^{-1})^{\beta}_{\mu} \Lambda^{\mu}_{\alpha} F^{\alpha}_{\beta}$$

$$= g^{\beta}_{\alpha} F^{\alpha}_{\beta} = F^{\beta}_{\beta}$$

→ transforms as scalar

$A^{\mu}_{\nu\alpha\beta}$: tensor of 4th rank

$A^{\mu}_{\mu\alpha\beta} \rightarrow$ transforms as covariant tensor of 2nd rank

Definitions of differential operators

4- gradient

$$\begin{aligned} \text{a) } \partial_\mu &= \frac{\partial}{\partial x^\mu} = \left(\frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla} \right) && \text{covariant vector} \\ \text{b) } \partial^\mu &= \frac{\partial}{\partial x_\mu} = \left(\frac{1}{c} \frac{\partial}{\partial t}, -\vec{\nabla} \right)^T && \text{contravariant vector} \end{aligned}$$

4- divergence

$$\partial^\mu a_\mu = \partial_\mu a^\mu = \frac{1}{c} \frac{\partial a^0}{\partial t} + \vec{\nabla} \cdot \vec{a} \quad (\text{scalar})$$

D'Alembert operator (box)

$$\square = \partial_\mu \partial^\mu = \partial^\mu \partial_\mu = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \Delta \quad (\text{scalar})$$

differentials

$$\begin{aligned} dx^\mu &= (c dt, d\vec{r})^T && (4\text{-vector}) \\ ds^2 &= d(x_\mu x^\mu) = c^2 (dt)^2 - (d\vec{r})^2 && \text{scalar} \end{aligned}$$

proper time

$$\begin{aligned} d\tau^2 &= \frac{1}{c^2} ds^2 && \text{defined for time-like} \\ &&& ds^2 > 0 \\ d\tau &= \frac{1}{c} \sqrt{ds^2} \end{aligned}$$

consider: object described

1) in a reference system where it rests

$$d\vec{r} = 0, \quad (d\vec{r})^2 = 0$$

$$ds^2 = c^2 (dt)^2$$

$$\text{proper time} \quad d\tau = \frac{1}{c} \sqrt{ds^2} = \frac{1}{c} \sqrt{c^2 (dt)^2} = dt$$

2) in an arbitrary reference frame

$$\begin{aligned}
 c^2(d\tau)^2 &= c^2(dt^2) - (d\vec{r})^2 = c^2(dt)^2 \left[1 - \frac{(d\vec{r})^2}{c^2(dt)^2} \right] \\
 &= c^2(dt)^2 \left[1 - \frac{1}{c^2} \left(\frac{d\vec{r}}{dt} \right)^2 \right] \\
 \frac{d\vec{r}}{dt} = \vec{v} &\quad \rightarrow \quad = c^2(dt)^2 \left[1 - \frac{v^2}{c^2} \right] = c^2(dt)^2 [1 - \beta^2] \\
 &= c^2(dt)^2 \frac{1}{\gamma^2}
 \end{aligned}$$

$$\Rightarrow d\tau = \frac{dt}{\gamma} \quad \Leftrightarrow dt = \gamma d\tau \quad (\text{time dilatation})$$

Physical quantities

1) 4-velocity

$$u^\mu = \frac{dx^\mu}{d\tau} \quad \text{is a contravariant vector}$$

since x^μ is one and τ is scalar

$$\begin{aligned}
 u^\mu &= \frac{d}{d\tau} (x^\mu) = \frac{d}{d\tau} (ct, \vec{r})^\top \\
 &= \gamma \frac{d}{dt} (ct, \vec{r}) = \gamma (c, \vec{v})
 \end{aligned}$$

construct Lorentz scalar:

$$\begin{aligned}
 u^\mu u_\mu &= \gamma^2 (c^2 - \vec{v}^2) = \frac{1}{1 - \beta^2} \left(c^2 - \frac{\beta^2}{c^2} \right) \\
 &= \frac{1}{1 - \beta^2} (1 - \beta^2) c^2 = c^2
 \end{aligned}$$

2) 4-momentum

$$p^\mu = m u^\mu \quad m: \text{mass, Lorentz scalar}$$

(contravariant vector)

construct Lorentz scalar

$$p^\mu p_\mu = m^2 u^\mu u_\mu = m^2 c^2$$

a closer look at the components

$$p^\mu = (p^0, \vec{p})^T = m\gamma (c, \vec{v})^T \quad \begin{aligned} p^0 &= m\gamma c \\ \vec{p} &= m\gamma \vec{v} \end{aligned}$$

for $v \ll c$, we have $\gamma = 1 + O\left(\frac{v^2}{c^2}\right)$

$$\vec{p} = m\gamma \vec{v} \approx m\vec{v} = \vec{p}_{\text{nonrel}}$$

3) relativistic generalization of Newton's equation

4-force (Minkowski force)

$$K^\mu = \frac{d}{d\tau} p^\mu = m \frac{du^\mu}{d\tau}$$

a closer look at the components

$$K^\mu = (K^0, \vec{K}) = \left(\gamma \frac{dp^0}{dt}, \gamma \frac{d\vec{p}}{dt} \right)^T$$

$$K^0 = \gamma \frac{dp^0}{dt} = \gamma \frac{d}{dt} (m\gamma c) = m\gamma c \frac{d\gamma}{dt}$$

$$\vec{K} = m\gamma \frac{d(\gamma \vec{v})}{dt}$$

define a 3-vector force

$$\vec{F} = \frac{d\vec{p}}{dt} \quad \text{such that} \quad \vec{K} = \gamma \vec{F}$$

and we get the nonrelativistic limit for $v \ll c$

$$\vec{F} = \frac{d\vec{p}}{dt} = \frac{d}{dt} (\gamma \vec{p}_{\text{nonrel}}) \approx \frac{d\vec{p}_{\text{nonrel}}}{dt} = \vec{F}_{\text{nonrel}}$$

(Newton's equation)

4) Relativistic energy momentum relation

Question: How to interpret K^0 ?

construct Lorentz scalar product rule + metric tensor $g^{\mu\nu}$

$$K^\mu u_\mu = m \frac{du^\mu}{d\tau} u_\mu = \frac{m}{2} \frac{d}{d\tau} (u^\mu u_\mu)$$

$$= \frac{m}{2} \frac{d}{d\tau} (c^2) = 0$$

in other notation: $K \cdot U = 0$

4-force is "orthogonal" to 4-velocity

a closer look to components

$$0 = K^\mu u_\mu = K^0 \gamma c - \vec{K} \cdot (\gamma \vec{v})$$

$$= \gamma c \left(K^0 - \gamma \vec{F} \cdot \frac{\vec{v}}{c} \right)$$

read off: $K^0 = \gamma \frac{\vec{F} \cdot \vec{v}}{c}$

nonrelativistic mechanics $\vec{F} \cdot \vec{v} = -P$ (power)

we obtain: $K^0 = m \gamma c \frac{d\gamma}{dt}$

$$-P = \vec{F} \cdot \vec{v} = \frac{c}{\gamma} K^0 = m c^2 \frac{d\gamma}{dt} = \frac{d}{dt} (m \gamma c^2)$$

more on nonrelativistic mechanics:

$$P = -\dot{T} \quad (dW = -dT)$$

we can read off the relativistic generalization of the kinetic energy

$$T = m \gamma c^2, \quad \text{then} \quad -P = \vec{F} \cdot \vec{v} = \frac{dT}{dt} = \dot{T}$$