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(6)

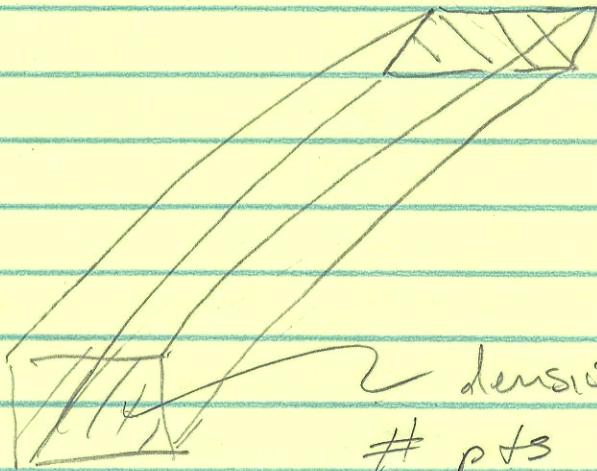
Announcements

Reading: Ch. 8

Hw 2 Sept. 19, Hw 1 Solns posted

Quiz remarks

Last time Liouville's Theorem



$$p = \frac{dN}{dq dp} \quad \text{Analogy to Fluid: } \frac{\partial S}{\partial E} + \nabla \cdot \vec{j} = 0$$

For inf. region

$$\frac{dp_k}{dq_k}$$

we calculated flux in - flux out

$$= \frac{\partial S}{\partial q_k} \dot{q}_k + p \frac{\partial \dot{q}_k}{\partial q_k} + \frac{\partial p}{\partial p_k} \dot{p}_k + p \frac{\partial \dot{p}_k}{\partial p_k}$$

① ② ③ ④

(1)

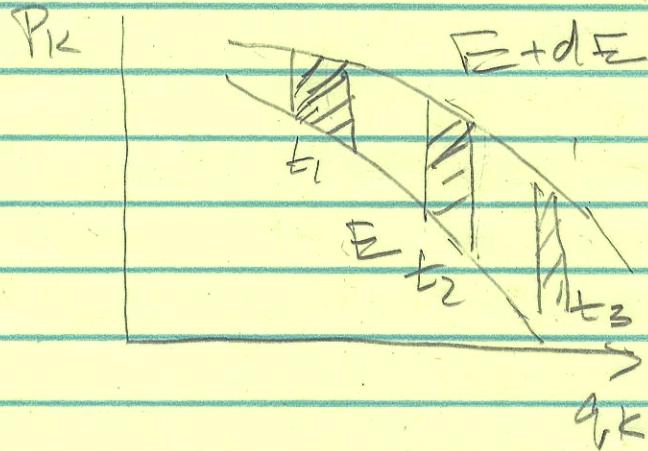
$$\textcircled{2} + \textcircled{4} \text{ cancel since } \dot{q}_k = \frac{\partial H}{\partial p_k}; p_k = -\frac{\partial H}{\partial q_k}$$

$$\Rightarrow \textcircled{2} + \textcircled{4} = \frac{\partial H}{\partial q_k \partial p_k} + \frac{\partial H}{\partial p_k \partial q_k} = 0$$

So our "continuity eqn" is (sum over k)

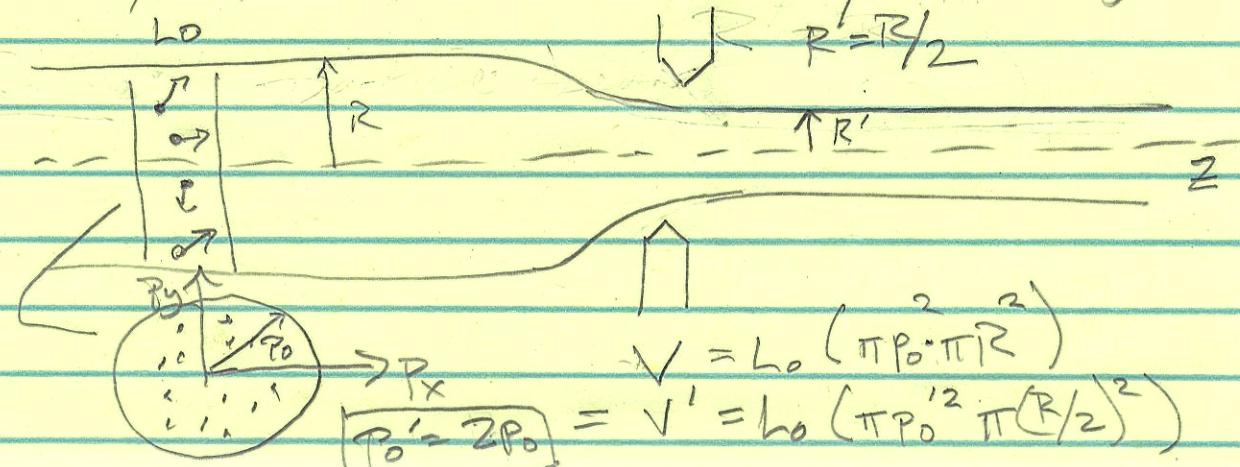
$$\frac{\partial \mathcal{P}}{\partial t} + \sum_k \left(\frac{\partial \mathcal{P}}{\partial q_k} \dot{q}_k + \frac{\partial \mathcal{P}}{\partial p_k} \dot{p}_k \right) = \frac{d\mathcal{P}}{dt} = 0$$

Draw 2 hypersurfaces in $2N$ -dim space



$\Delta p_k, \Delta q_k$ must change to preserve \mathcal{P}

Prob. 13.35
L₀ doesn't
change



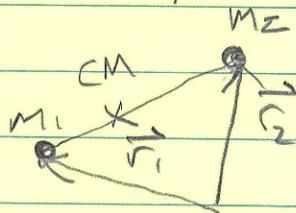
(2)

Ch. 8 Orbits w central forces

3D 2-body problem + reduction
to effective 1body, 1D problem.

(A)

CM + relative coords.



$$\text{CM position } \vec{R} = \frac{\vec{r}_1 + m_2 \vec{r}_2}{m_1 + m_2}$$

$$\text{total mass } M = m_1 + m_2 \quad \text{rel. position } \vec{r} = \vec{r}_1 - \vec{r}_2$$

$$\text{Invarit } \vec{r}_1 = \vec{R} + \frac{m_2}{M} \vec{r} \quad \vec{r}_2 = \vec{R} - \frac{m_1}{M} \vec{r}$$

$$\text{KE } \bar{T} = \frac{1}{2} M \dot{\vec{r}}_1^2 + \frac{1}{2} m \dot{\vec{r}}_2^2$$

$$= \frac{1}{2} M \dot{\vec{R}}^2 + \frac{1}{2} \left(\frac{m_1 m_2}{M} \right) \dot{\vec{r}}^2$$

CM motion

relative motion

$$\text{reduced mass } \mu = \frac{m_1 m_2}{M}$$

$$\mathcal{L} = \bar{T} - \mathcal{U} = \frac{1}{2} M \dot{\vec{R}}^2 + \left(\frac{1}{2} \mu \dot{\vec{r}}^2 - U(\vec{r}) \right)$$

 \mathcal{L}_{cm} \mathcal{L}_{rel}

Consider only those cases where
 $U(\vec{r}_1, \vec{r}_2) = U(r) \quad \begin{matrix} \text{Gravity} \\ \text{Coulomb, ...} \end{matrix}$

$$\vec{R} = (x, y, z) \quad i=x, y, z$$

(3)

$$\text{CM motion} \quad \frac{\partial L_{\text{CM}}}{\partial \dot{x}_i} = \frac{d}{dt} \frac{\partial L_{\text{CM}}}{\partial \ddot{x}_i} = M \ddot{x}_i = 0$$

$\Rightarrow P_{\text{CM}}$ conserved

$$\frac{d}{dt} (M \dot{x}_i) = 0$$

\vec{R} = "ignorable" L depends only on \vec{R}

$$\frac{\partial L_{\text{rel}}}{\partial x_i} = \frac{d}{dt} \frac{\partial L}{\partial \dot{x}_i}$$

$$-\frac{\partial U}{\partial x_i} = \frac{d}{dt} \frac{\partial L_{\text{rel}}}{\partial \dot{x}_i} = \mu \ddot{x}_i$$

\Rightarrow single particle w/ mass μ

* Now we've reduced the problem to an effective one particle, 3D one. However, we can go further!

$$\vec{L} = \vec{r}_1 \times \vec{p}_1 + \vec{r}_2 \times \vec{p}_2$$

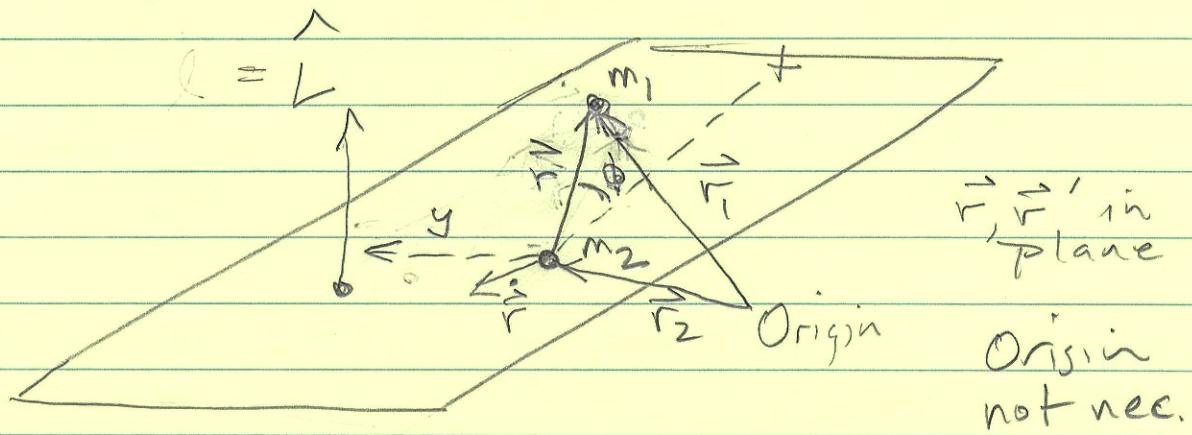
$$\text{In CM frame } \vec{R} = 0 \quad \vec{r}_1 = \frac{m_2}{M} \vec{r} \quad \vec{r}_2 = \frac{m_1}{M} \vec{r}$$

$$\Rightarrow \boxed{\vec{L} = \frac{m_1 m_2}{M} \vec{r} \times \vec{v} = \left(\vec{r} \times \mu \vec{v} \right)}$$

i.e. same form as our single particle, mass μ position \vec{r}

4

Since there is no external force, \vec{L} is conserved. In particular \vec{L} is same throughout motion. This $\Rightarrow \vec{r}, \vec{r}'$ are both $\perp \vec{L}$, i.e. they always lie in a plane.



Origin
not nec.

Choose coordinates in this plane r, ϕ

$$\mathcal{L} = \frac{1}{2} \mu (r^2 + r^2 \dot{\phi}^2) - u(r)$$

$\boxed{\phi}$ \mathcal{L} ind. of $\phi - \phi$ is ignorable too!

A.M. cons. : $\frac{d\mathcal{L}}{d\phi} = 0 = \frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \frac{d}{dt} (\mu r^2 \dot{\phi})$

$$\ell = L_z$$

check
this!

$$\vec{r} = r(\cos\phi, \sin\phi)$$

$$\vec{r}' = \dot{r}(\cos\phi, \sin\phi) + r(-\sin\phi, \cos\phi)$$

use: $\vec{r} \times \vec{r}' = r^2 \dot{\phi}$