Final Exam Solutions–Spring 2023

Problem 1

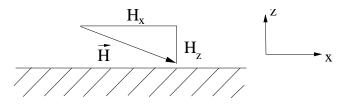
No "right answer"!.

Problem 2- one-phonon neutron scattering

Ashcroft-Mermin Problem 24-2 Allowed emitted phonon solutions follows from Figure 24. 6 but taking $\omega(\mathbf{k}) \rightarrow -\omega(\mathbf{k})$.

Problem 3 - Superconducting sphere

1. Outside sphere must have $\vec{\nabla} \cdot \vec{H} = 0$, $\vec{\nabla} \times \vec{H} = 0$ Near surface, picture like this locally



Since on these scales system is symmetric, \vec{H} can only depend on z. Therefore

$$0 = \vec{\nabla} \cdot \vec{H} = \nabla_z H_z(z) + \nabla_x H_x(z) = \nabla_z H_z$$

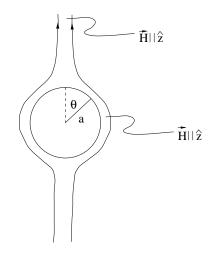
So H_z is at most a constant ind. of z everywhere. This implies $\vec{\nabla} \times \vec{H} = 0$ inside sphere too \Rightarrow no currents flowing \Rightarrow trivial solution.

The above shows that we must have $\vec{H}\parallel$ surface of sphere everywhere for Meissner effect.

2. Look for soln. to $\vec{\nabla} \cdot \vec{H} = 0$, $\vec{\nabla} \times \vec{H} = 0$ outside sphere with BC. $H_{\perp}|_{r=a} = 0$, $\vec{H} = H_0 \hat{z}$ for r >> a, a = radius of sphere.

1

Guess solution with form



In exterior this is standard magnetostatics problem, with sphere of uniform magnetization. Magnetostatic potential is dipolar:

$$\Phi_{H} = \frac{4}{3}\pi a^{3} M \frac{\cos \theta}{r^{2}}$$
$$\vec{H} = \vec{H}_{0} - \vec{\nabla} \frac{4}{3}\pi a^{3} M \frac{\cos \theta}{r^{2}} = \vec{H}_{0} + \frac{4}{3}\frac{\pi a^{3} M}{r^{3}}(2\cos \theta \,\hat{r} + \sin \theta \,\hat{\theta})$$

Require $H_r|_{r=a} = 0$ B.C.

$$=H_0\cos\theta+\frac{4\pi}{3}M\,2\cos\theta=0$$

effective mag. "M" = $\frac{3}{8\pi}H_0$

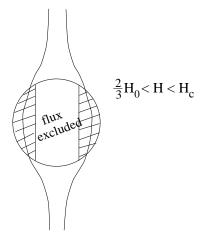
$$\vec{H} = H_0 \left(\hat{z} + \frac{a^3}{2r^3} (2\cos\theta \,\hat{r} + \sin\theta \,\hat{\theta}) \right)$$

Component \parallel to surface is

$$H_{\theta}|_{r=a} = H_0 \sin \theta (1 + \frac{1}{2}) = \frac{3}{2} H_0 \sin \theta$$

 $(z = \cos \theta \, \hat{r}, \sin \theta \, \hat{\theta}).$

So when $H_0 = \frac{2}{3}H_c$, field at equator $H(\theta = \pi/2, r = a) = H_c \Rightarrow$ SC at that point is supressed. Energetically favorable for part of sample near equator to become normal:



(N.B. when $H_0 > H_c$, excluding flux is no longer energetically favorable, and solution above no longer applies).

Problem 4 - Density of states of p-wave superconductors.

1. Suppose that a superconductor is described by an order parameter whose momentum dependence is $\Delta_{\mathbf{k}} = \Delta_0 f(\hat{k})$, where $f(\hat{k})$ is either $\sin \theta$ (axial state) or $\cos \theta$ (polar state). Note the two f functions are $\ell = 1$ spherical harmonics or linear combinations thereof, so to satisfy the Pauli principle these must represent spin triplet (S = 1) pair states. Nevertheless the quasiparticle spectrum may be taken to be independent of spin, $E_{\mathbf{k}} = \sqrt{\xi_{\mathbf{k}}^2 + \Delta_{\mathbf{k}}^2}$.

Use the fact that, for any given angle θ of **k** on the Fermi surface, the quasiparticle excitations have the same form as that for an *s*-wave superconductor, but with gap $\Delta(\theta)$ to find an expression for the density of states for the two types of order parameters.

In class we showed that the density of quasiparticle states of a superconductor is $N(\omega) = N_0 \operatorname{Re} \omega / \sqrt{\omega^2 - \Delta^2}$. We are told this holds now for each angle,

$$N(\omega; \theta) = N_0 \text{ Re } \frac{\omega}{\sqrt{\omega^2 - \Delta^2(\theta)}}$$

The density of states is now just N_0 ReI.

So for the total density of states, we need the two integrals

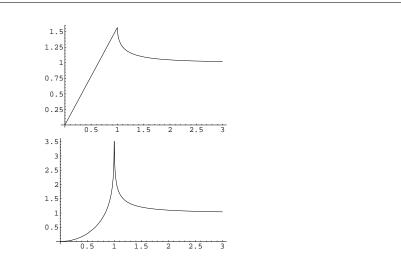
$$I_{pol} = \int_0^{\pi/2} d\theta \sin \theta \frac{\omega}{\sqrt{\omega^2 - \Delta_0^2 \cos^2 \theta}} = \frac{\omega}{\Delta_0} \cot^{-1} \sqrt{\omega^2 - \Delta_0^2}$$
$$\int_0^{\Pi/2} \omega \sqrt{\omega^2 - \Delta_0^2} d\theta \sin \theta \frac{\omega}{\sqrt{\omega^2 - \Delta_0^2 \cos^2 \theta}} = \frac{\omega}{\Delta_0} \cot^{-1} \sqrt{\omega^2 - \Delta_0^2}$$

 polar

$$I_{ax} = \int_0^{\Pi/2} d\theta \sin \theta \frac{\omega}{\sqrt{\omega^2 - \Delta_0^2 \sin^2 \theta}} = \omega \log \sqrt{\frac{\omega + \Delta_0}{\omega - \Delta_0}}$$

axial,

2. Plot as a function of ω the full frequency dependence of the density of states in the two cases.



Density of states $N(\omega)/N_0$ vs. ω for polar (top) and axial (bottom) *p*-wave states. Note in both cases as $\omega \to \infty$ the functions approach 1, i.e. the normal state DOS.

3. Estimate the low- ω asymptotic behavior in both cases.

Expanding the functions $I_{ax,pol}$ for small ω gives $I_{pol} \simeq \frac{\omega}{\Delta_0} \left(\pi/2 + i \log \omega^2 / 4\Delta_0^2 \right)$ $I_{ax} = \frac{\omega}{\Delta_0} \left(\frac{\omega}{\Delta_0} - i\frac{\pi}{2} \right)$

so the desired asymptotic forms are $N_{pol} \simeq \pi \omega/(2\Delta_0)$ and $N_{ax} \simeq \omega^2/\Delta_0^2$. These two power laws are reflected directly in the temperature dependence of the specific heat of the two states.

Problem 5: 1D Ising Model.

Consider a chain of N sites governed by the Hamiltonian

$$H_I = J \sum_{i=1}^{N-1} S_i^z S_{i+1}^z.$$
 (1)

Consider a chain of N sites governed by the Hamiltonian

$$H_I = J \sum_{i=1}^{N-1} S_i^z S_{i+1}^z.$$
 (2)

Show that the partition function is given by

$$Z = \sum_{S_1^z = \pm \frac{1}{2}} \dots \sum_{S_N^z = \pm \frac{1}{2}} \exp\left(-\beta J \sum_{i=1}^{N-1} S_i^z S_{i+1}^z\right) = 2 \left(2 \cosh[\beta J/4]\right)^{N-1} (3)$$

Start from the end of the chain and show that $\sum_{S_N^z = \pm \frac{1}{2}} \exp\left(-\beta J S_{N-1}^z S_N^z\right) = 2 \cosh(\beta J/4)$ independent of S_{N-1}^z . Thus $Z_N = 2 \cosh(\beta J/4) Z_{N-1}$. Iterate the procedure, and remember to be careful at the other end of the chain to get the result.

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The thermodynamic limit is defined by $N \to \infty$.

Partition function is

$$Z = \sum_{S_1^z = \pm 1/2} \sum_{S_2^z = \pm 1/2} \cdots \sum_{S_N^z = \pm 1/2} \exp\left[-\beta J \sum_{i=1}^{N-1} S_i^z S_{i+1}^z\right]$$

$$= \sum_{S_2^z = \pm 1/2} \cdots \sum_{S_N^z = \pm 1/2} \left(e^{\beta \frac{J}{2} S_2^z} + e^{-\beta \frac{J}{2} S_2^z}\right) e^{-\beta J S_2^z S_3^z} \cdots e^{-\beta J S_{N-1}^z S_N^z}$$

$$= \sum_{S_3^z = \pm 1/2} \cdots \sum_{S_N^z = \pm 1/2} \left[2 \cosh(\beta J/4) e^{-\beta \frac{J}{2} S_3^z} + 2 \cosh(\beta J/4) e^{\beta \frac{J}{2} S_3^z}\right] e^{-\beta J S_3^z S_4^z} \cdots e^{-\beta J S_{N-1}^z S_N^z}$$

$$= \left[2 \cosh\beta J/4\right]^2 \sum_{S_4^z = \pm 1/2} \cdots \sum_{S_N^z = \pm 1/2} \left(e^{\beta \frac{J}{2} S_4^z} + e^{-\beta \frac{J}{2} S_4^z}\right) e^{-\beta J S_4^z S_5^z} \cdots e^{-\beta J S_{N-1}^z S_N^z}$$

$$\vdots$$

$$= 2\left[2 \cosh\beta J/4\right]^{N-1}$$
(4)

Show that in the thermodynamic limit the free energy is given by

$$F = -NkT\ln(2\cosh(\beta J/4))$$

$$F = -T \ln Z = -T \ln \left[2(2 \cosh(\beta J/4))^{N-1} \right] = -T \ln 2 - T(N-1) \ln \left[2 \cosh(\beta J/4) \right] = -T \left[(\ln 2 - \ln(2 \cosh(\beta J/4))) - TN \ln \left[2 \cosh(\beta J/4) \right] \right]$$

(5)

 $\rightarrow_{N \to \infty} -TN \ln \left[2 \cosh(\beta J/4) \right]$ (6)

Let us calculate the magnetization per site $M = -\frac{1}{N} \frac{\partial F}{\partial H}|_{H=0}$. We'll need to include the coupling to the magnetic field. Thus we need to know the free energy in the presence of a magnetic field H where

$$H_I = J \sum_{i=1}^{N} S_i^z S_{i+1}^z - H \sum_{i=1}^{N} S_i^z.$$
 (7)

For convenience we now sum from i = 1, ..., N assuming periodic boundary conditions meaning that N + 1 = 1

Convince yourself that the partition function can be written as

$$Z = \sum_{S_i^z = \pm \frac{1}{2}} \prod_{i=1}^N \exp\left(-\beta \left[JS_i^z S_{i+1}^z - \frac{H}{2}(S_i^z + S_{i+1}^z)\right]\right)$$
(8)

Consider the Ising chain with N sites in magnetic field H with boundary conditions $S_{N+1} = S_1$.

$$\begin{aligned} \mathcal{H} &= J \sum_{i=1}^{N} S_{i}^{z} S_{i+1}^{z} + H \sum_{i=1}^{N} S_{i}^{z} \\ Z &= \sum_{S_{1}^{z} = \pm 1/2} \sum_{S_{2}^{z} = \pm 1/2} \cdots \sum_{S_{N}^{z} = \pm 1/2} \exp\left[-\beta J \sum_{i=1}^{N-1} S_{i}^{z} S_{i+1}^{z} - \beta J S_{N}^{z} S_{1}^{z} - H \sum_{i=1}^{N} S_{i}^{z}\right] \\ &= \sum_{S_{1}^{z} = \pm 1/2} \cdots \sum_{S_{N}^{z} = \pm 1/2} \exp\left[-\beta J \left(S_{1}^{z} S_{2}^{z} + S_{2}^{z} S_{3}^{z} + \dots S_{N}^{z} S_{1}^{z}\right) - \beta H \left(S_{1}^{z} + S_{2}^{z} + \dots + S_{N}^{z}\right)\right] \\ &= \sum_{S_{1}^{z} = \pm 1/2} \cdots \sum_{S_{N}^{z} = \pm 1/2} \exp\left(-\beta J S_{1}^{z} S_{2}^{z} - \beta H S_{1}^{z}\right) \exp\left(J S_{2}^{z} S_{3}^{z} - \beta H S_{2}^{z}\right) \cdots \exp\left(-\beta J S_{N}^{z} S_{1}^{z} - \beta H S_{N}^{z}\right) \\ &= \sum_{S_{1}^{z} = \pm 1/2} \cdots \sum_{S_{N}^{z} = \pm 1/2} \exp\left(-\beta J S_{1}^{z} S_{2}^{z} - \beta \frac{H}{2} \left(S_{1}^{z} + S_{2}^{z}\right)\right) \cdots \exp\left(-\beta J S_{N}^{z} S_{1}^{z} - \beta \frac{H}{2} \left(S_{N}^{z} + S_{1}^{z}\right)\right) \\ &= \sum_{S_{1}^{z} = \pm 1/2} \cdots \sum_{S_{N}^{z} = \pm 1/2} \prod_{i=1}^{N} \exp\left(-\beta J S_{1}^{z} S_{2}^{z} - \beta \frac{H}{2} \left(S_{1}^{z} + S_{2}^{z}\right)\right) \cdots \exp\left(-\beta J S_{N}^{z} S_{1}^{z} - \beta \frac{H}{2} \left(S_{N}^{z} + S_{1}^{z}\right)\right) \end{aligned}$$
(9)

This partition function can be rewritten in terms of a $\mathit{transfer \ matrix}\ P$ given by

$$P = \begin{pmatrix} P_{11} & P_{1-1} \\ P_{-11} & P_{-1-1} \end{pmatrix}$$
(10)

where

$$P_{11} = \exp(-\beta(J+H)),$$
(11)

$$P_{-1-1} = \exp(-\beta(J - H)), \tag{12}$$

$$P_{1-1} = P_{-11} = \exp(\beta J), \tag{13}$$

with $\tilde{J} = J/4$ and $\tilde{H} = H/2$.

Show that
$$Z = \text{Tr}P^N.$$
 (14)

First note, e.g. that 1 and -1 stand for spin up or down, so, e.g. $\langle \uparrow | P | \uparrow \rangle = \exp(-J/4 - H/2)$, which is just the first exponential factor in the framed equation (8) above, for the term with both spins up. Check all the matrix elements $\uparrow \downarrow$, $\downarrow \downarrow$, and $\downarrow \uparrow$, to see that the expression derived above for Z is just

$$Z = \sum_{S_1^z} \cdots \sum_{S_N^z} \langle S_1^z | P | S_2^z \rangle \langle S_2^z | P | S_3^z \rangle \cdots \langle S_{N-1}^z | P | S_N^z \rangle \langle S_N^z | P | S_1^z \rangle$$
(15)

But each factor of $\sum_{S_i^z} |S_i^z\rangle \langle S_i^z| = 1$, so one can remove all the intermediate bras and kets except 1:

$$Z = \sum_{S_1^z} \langle S_1^z | P \cdot P \cdots P | S_1^z \rangle = \sum_{S_1^z} \langle S_1^z | P^N | S_1^z \rangle = \text{Tr} P^N$$
(16)

The idea is now to use that the trace Tr is basis independent. Thus we can diagonalise P and use this to obtain $\text{Tr}P^N$.

Find the eigenvalues λ_1 and λ_2 of P. Once you have them $Z = \text{Tr}P^N = \lambda_1^N + \lambda_2^N$.

Find eigenvalues λ_1 and λ_2 of a 2D matrix:

$$\begin{vmatrix} e^{-\beta(\tilde{J}+\tilde{H})} - \lambda & e^{\beta\tilde{J}} \\ e^{\beta\tilde{J}} & e^{-\beta(\tilde{J}-\tilde{H})} - \lambda \end{vmatrix} = 0$$
(17)

Solve

$$0 = \lambda^2 - 2e^{-\beta J} \cosh(\beta \tilde{H})\lambda - 2\sinh 2\beta \tilde{J}$$
(18)

$$(19)$$

$$\lambda_{\pm} = e^{-\beta \tilde{J}} \left[\cosh(\beta \tilde{H}) \pm \sqrt{\sinh^2(\beta \tilde{H}) + e^{4\beta \tilde{J}}} \right]$$
(20)

The partition function is now given by the Tr of P^N is now $\lambda^N_+ + \lambda^N_-$.

Use the fact that $\ln(\lambda_1^N + \lambda_2^N) = N \ln \lambda_1 + \ln \left(1 + \left[\frac{\lambda_2}{\lambda_1}\right]^N\right)$ to show that in the thermodynamic limit $F = -NkT \ln \left[e^{-\beta \tilde{J}} \cosh \beta \tilde{H} + \sqrt{e^{-2\beta \tilde{J}} \sinh^2 \beta \tilde{H} + e^{2\beta \tilde{J}}}\right]$ (21)

$$F = -T \ln Z = -T \ln (\lambda_{+}^{N} + \lambda_{-}^{N})$$

$$= -TN \ln \lambda_{+} - T \ln \left(1 + \left(\frac{\lambda_{-}}{\lambda_{+}} \right)^{N} \right)$$
(22)

Now note that $\lambda_{-}/\lambda_{+} < 1$, so in the thermodynamic limit $N \to \infty$, this vanishes.

$$F \xrightarrow{N \to \infty} -TN \ln \lambda_+ \tag{23}$$

For H = 0 this result reduces to Eq. (5) as it should. Now we are finally ready to obtain the magnetisation $M(\tilde{H})$ per site, i.e. $M(\tilde{H}) = -\frac{1}{N} \frac{\partial F}{\partial \tilde{H}}$.

Show that	
$M(\tilde{H}) = \frac{\sinh\beta\tilde{H}}{\sqrt{\sinh^2\beta\tilde{H} + e^{4\beta\tilde{J}}}}.$	(24)

$$M(\tilde{H}) = \frac{\partial F}{\partial \beta \tilde{H}} = \frac{\sinh \beta \tilde{H}}{\sqrt{\sinh^2 \beta \tilde{H} + e^{4\beta \tilde{J}}}}$$
(25)